

Dark Matter at Collider



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Discovery Time...

Dark Matter: we need new particles to explain the content of the universe

Standard Model: we need new physics

Supersymmetry solves both problems!

The super-partners are distributed around 100 GeV to a few TeV

LHC: directly probes TeV scale

Future results from PLANCK, direct and indirect detection, rare decays etc. experiments in tandem with the LHC will confirm a model

**This talk: How accurately we can calculate dark matter density?
Can we establish the existence particles responsible for 23%?**

So Far...

- Recent Higgs search results from Atlas and CMS indicate that Higgs mass (if it is Higgs) ~ 125 GeV
 - in the tight MSSM window: 115-135 GeV
- squark mass (first generation) \sim gluino mass ≥ 1.4 TeV
- For heavy squark mass, gluino mass is ≥ 900 GeV
- stop (squark) produced from gluinos, stop mass ≥ 700 GeV
- stop (squark) produced directly, stop mass ≥ 450 GeV
- Selectron/Smuon between 85 and 195 GeV for a 20 GeV neutralino are excluded at 95% confidence
- Chargino masses between 110 and 340 GeV are excluded at 95% CL for a neutralino of 10 GeV for Chargino decaying into e/μ

Dark Matter Content and SUSY...

Dark Matter content calculation:

Annihilation of lightest neutralinos \rightarrow SM particles

Annihilation diagrams: mostly non-colored particles, e.g., sleptons, staus, charginos, neutralinos, etc.

How to produce these non-colored particles at the LHC?

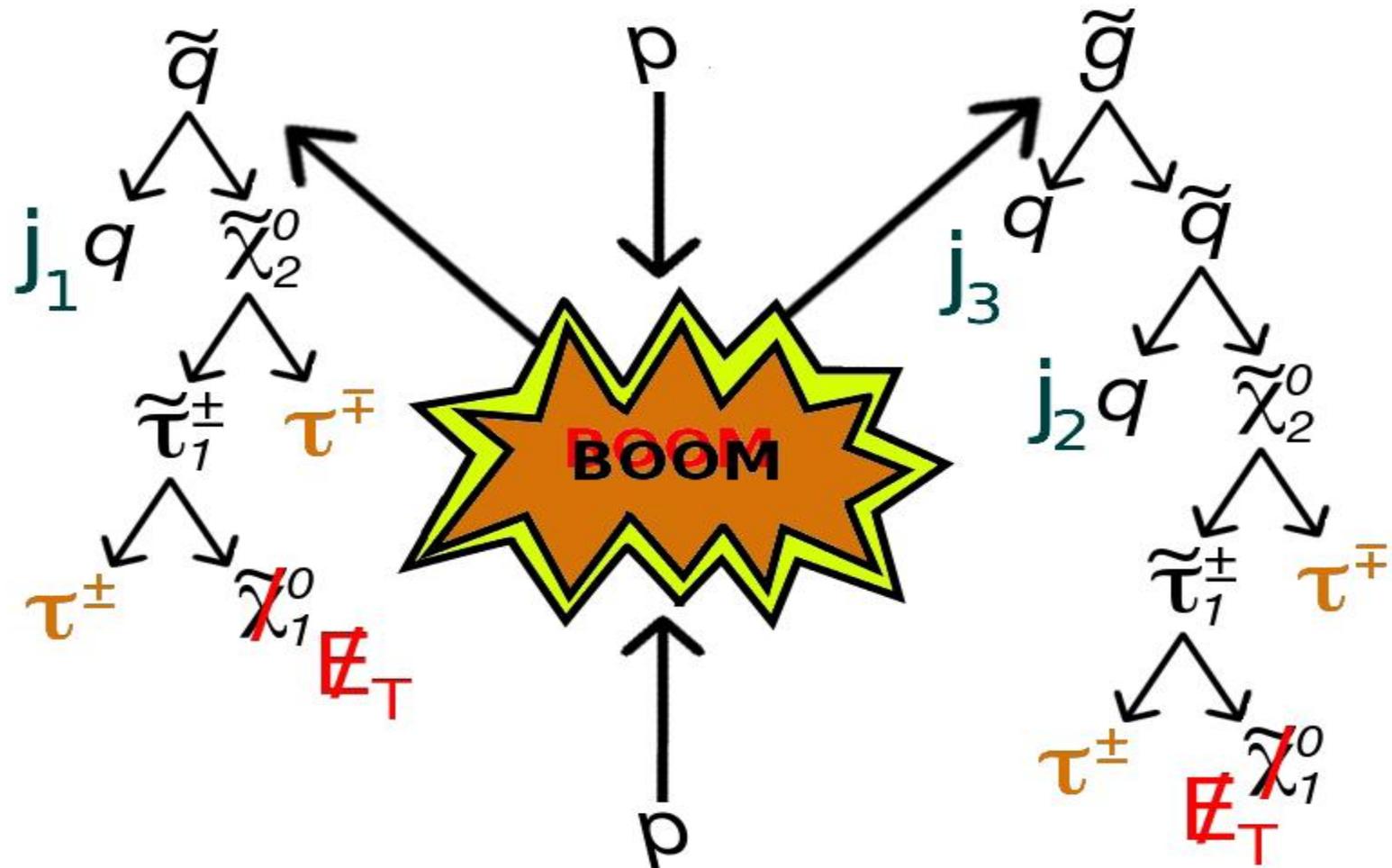
In this talk:

1. Cascade decays of squarks and gluinos

2. Vector Boson fusion

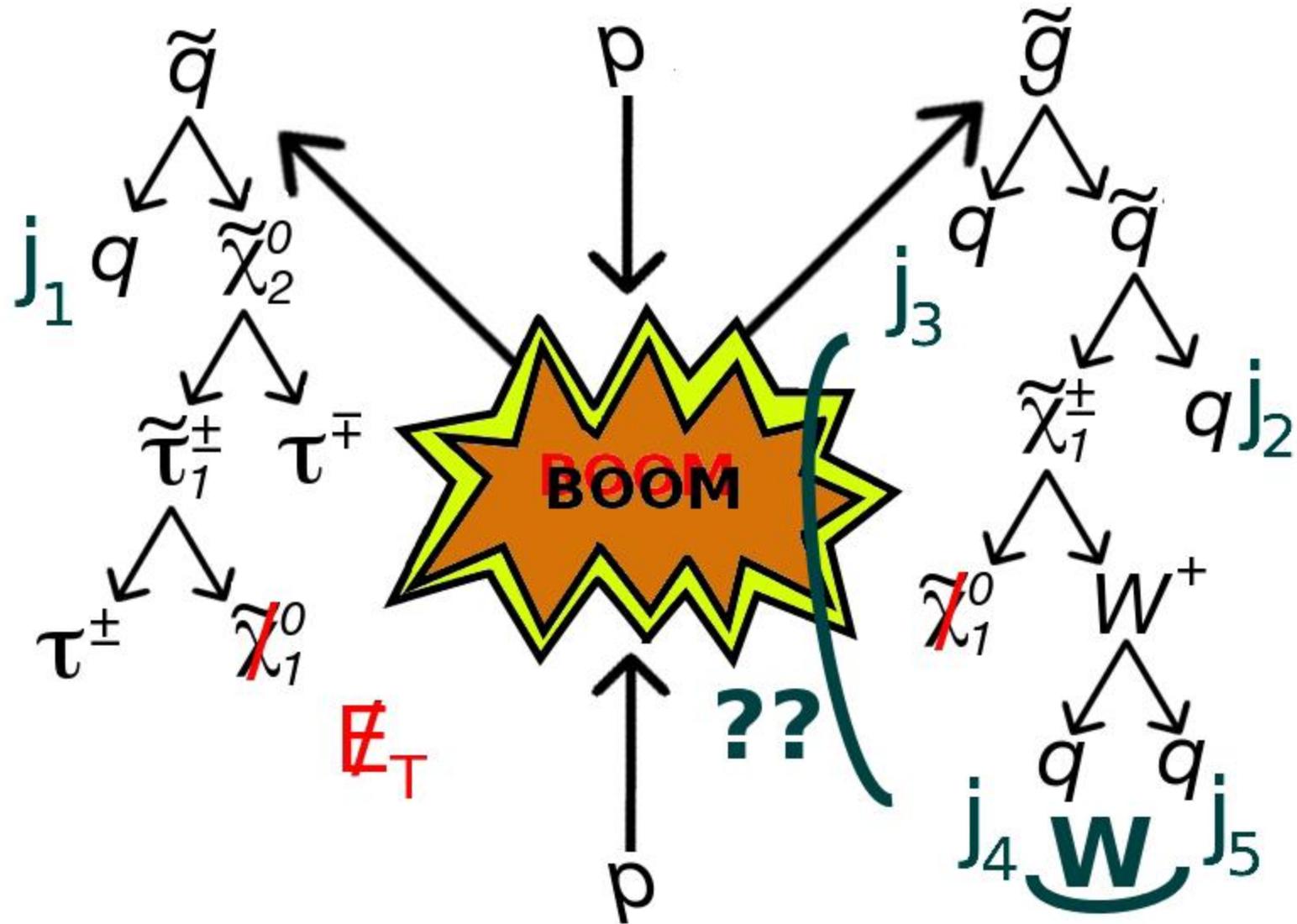
SUSY Particles via Cascade

Masses of particles are needed to calculate the DM content
But can we determine them?



SUSY Cascade @ LHC Dilemma

Can we determine all the masses in the diagram?



DM at the LHC

Goal:

Final states \rightarrow Masses \rightarrow Model Parameters

\rightarrow Calculate dark matter density

$$\tilde{Q} \rightarrow q + l + \tilde{\chi}_1^0$$

$$\tilde{L} \rightarrow l + \tilde{\chi}_1^0$$

$$\tilde{\chi}_{2,3,4}^0 \rightarrow Z, h, \bar{l}l + \tilde{\chi}_1^0 \quad \text{etc.}$$

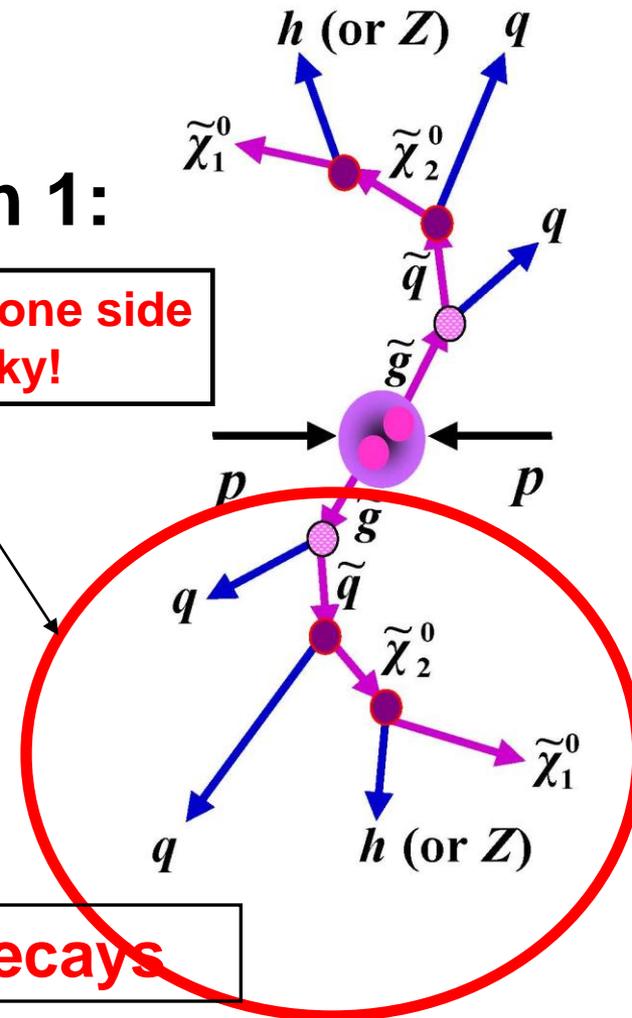
We may not be able to solve for masses of all the sparticles from a model

Problem 2:

Not all the sparticles appear in cascade decays

Problem 1:

Identifying one side is very tricky!



Solving for the MSSM : Very difficult

DM at the LHC via Cascade

Solutions:

Prob.1. Identifying one side is very tricky!

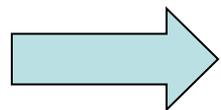
→ We develop new strategies: BEST, apply OS-LS

Prob. 2. Not all the sparticles appear in cascade decays

→ We can use simpler models to understand the cascades and solve for the model parameters

DM at the LHC via cascade

We can use simpler models to understand the cascades and solve for the model parameters



Calculate the Dark Matter content

The best strategy:

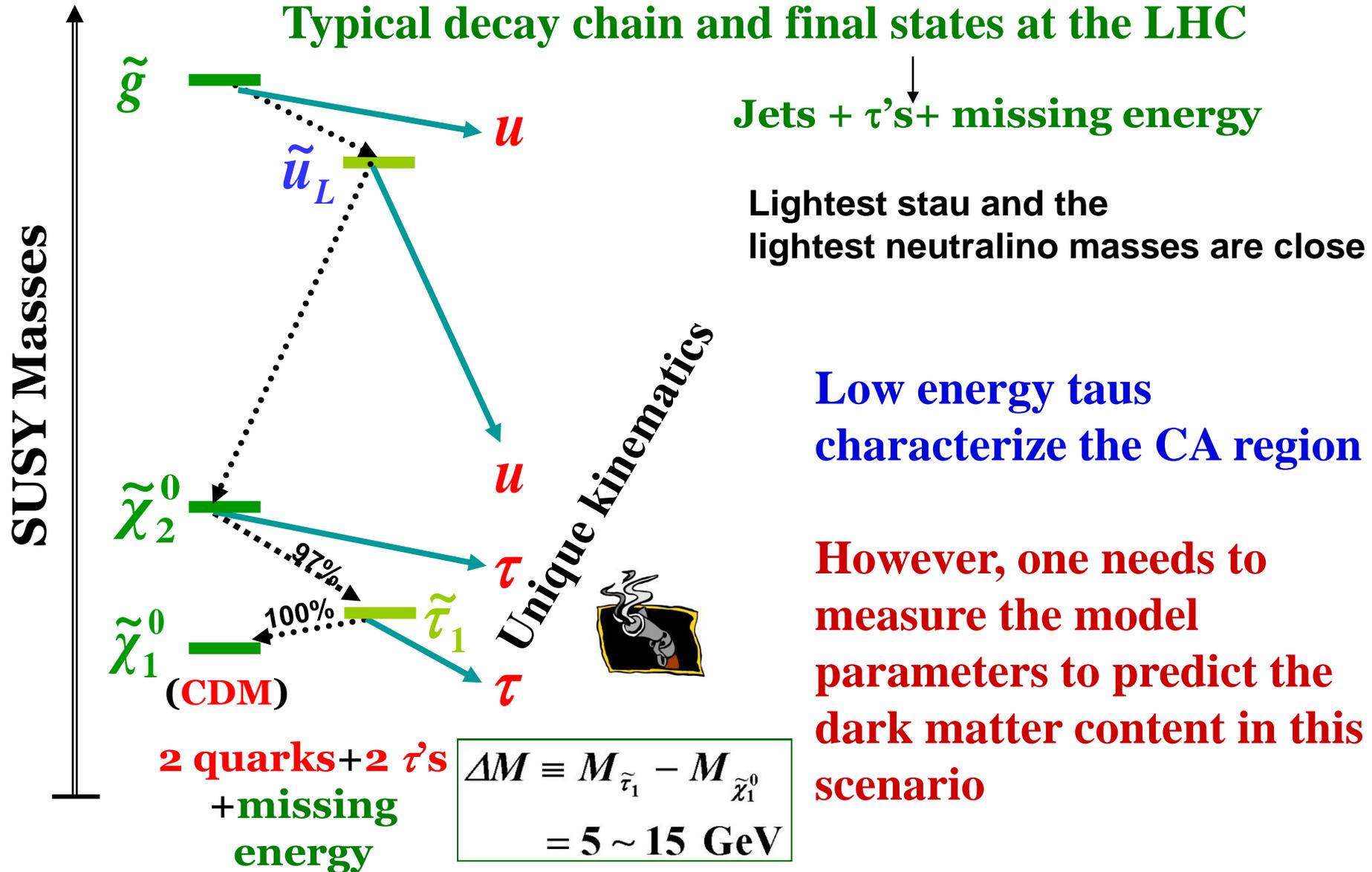
**Solve for the minimal model: mSUGRA/CMSSM →
4 parameters + sign: m_0 , $m_{1/2}$, A_0 , $\tan\beta$ and $\text{Sign}(\mu)$**

The cascades can be understood in a simpler way [hopefully!]

Also test:

**Models with more parameters or with different features, e.g.,
Next to minimal model (Higgs non-universality),
Gaugino Non-universality (Mirage Mediation model) etc...**

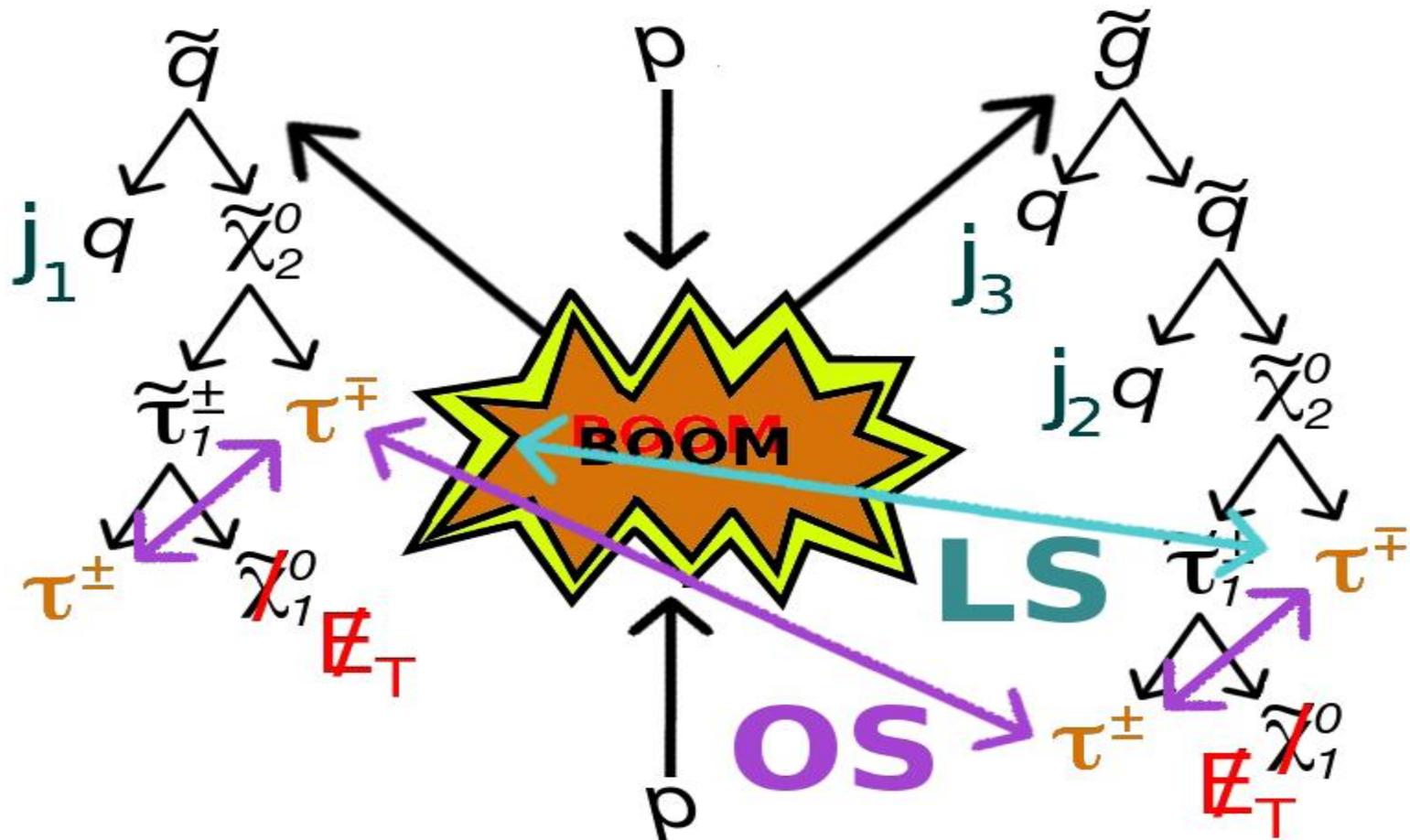
Case 1: Coannihilation Region



SUSY at the LHC Dilemma...

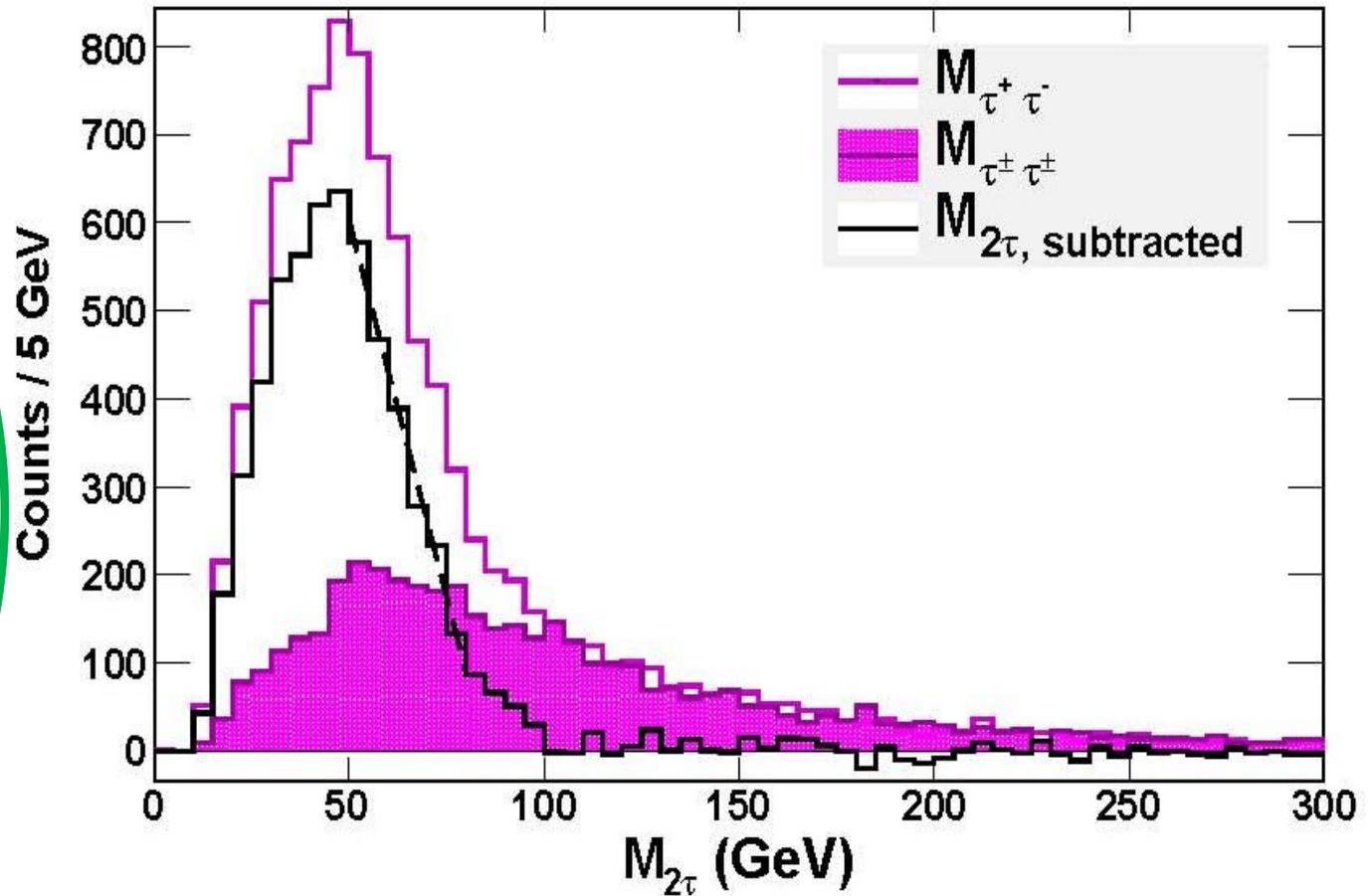
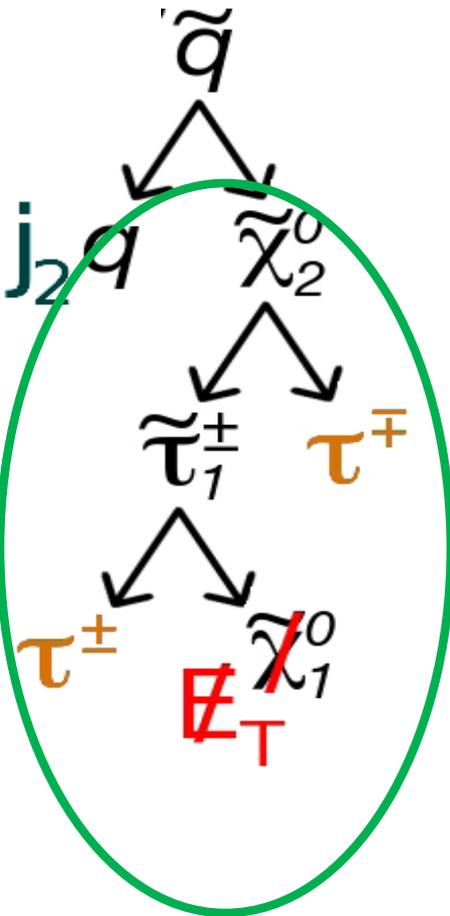
The stau and the lightest neutralino masses are needed to establish coannihilation region

But **identifying one side is tricky**



SUSY at the LHC Dilemma...

OS-LS Subtraction allows one side reconstruction

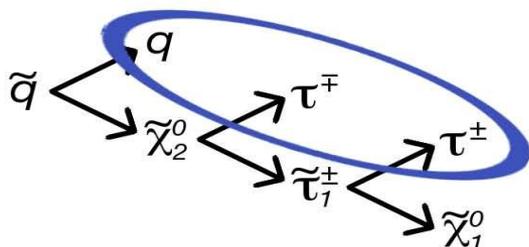


$$M_{\tau\tau}^{\text{end}} = M_{\tau\tau}^{\text{max}} = m_{\tilde{\chi}_2^0} \sqrt{\left(1 - \frac{m_{\tilde{\tau}_1}^2}{m_{\tilde{\chi}_2^0}^2}\right) \left(1 - \frac{m_{\tilde{\chi}_1^0}^2}{m_{\tilde{\tau}_1}^2}\right)} \rightarrow f(m_{\tilde{\tau}_1}, m_{\tilde{\chi}_2^0}, m_{\tilde{\chi}_1^0})$$

Extracting One side: $j\tau\tau$

➔ OS-LS selection of ditaus selects $\tilde{\chi}_2^0$, but if we need to reconstruct the entire side

➔ We use the following subtraction scheme:



The OS-LS τ pair has momentum related to the momentum of this **Same Event Jet**.

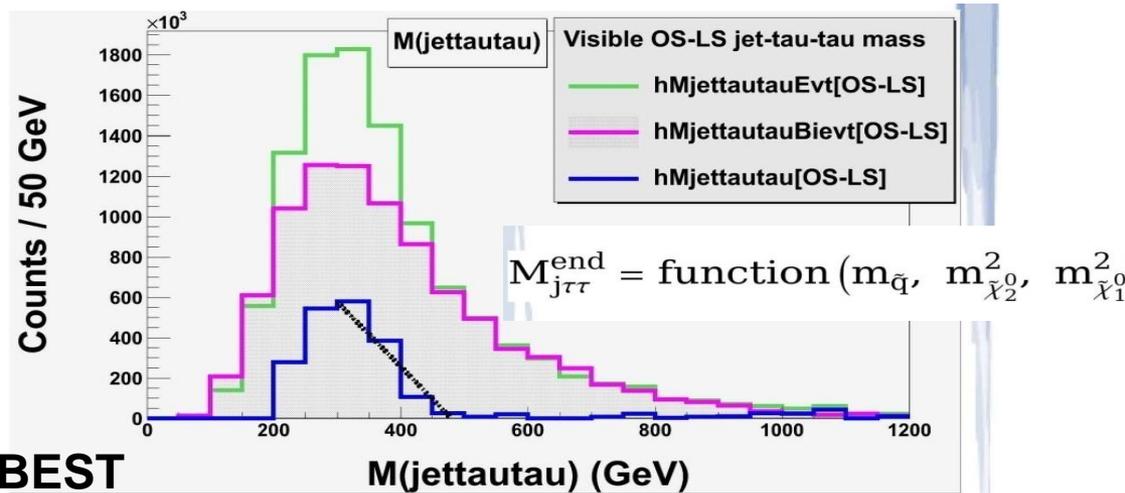
We collect all \blacksquare Jet pairs: get related pairs plus random pairs.

Using **Jets from Previous Events**: get only random pairs.

Normalize and perform the **Same Jet - Previous Jet** subtraction:

- Random pairs will cancel.
- Only the related pairs remain.

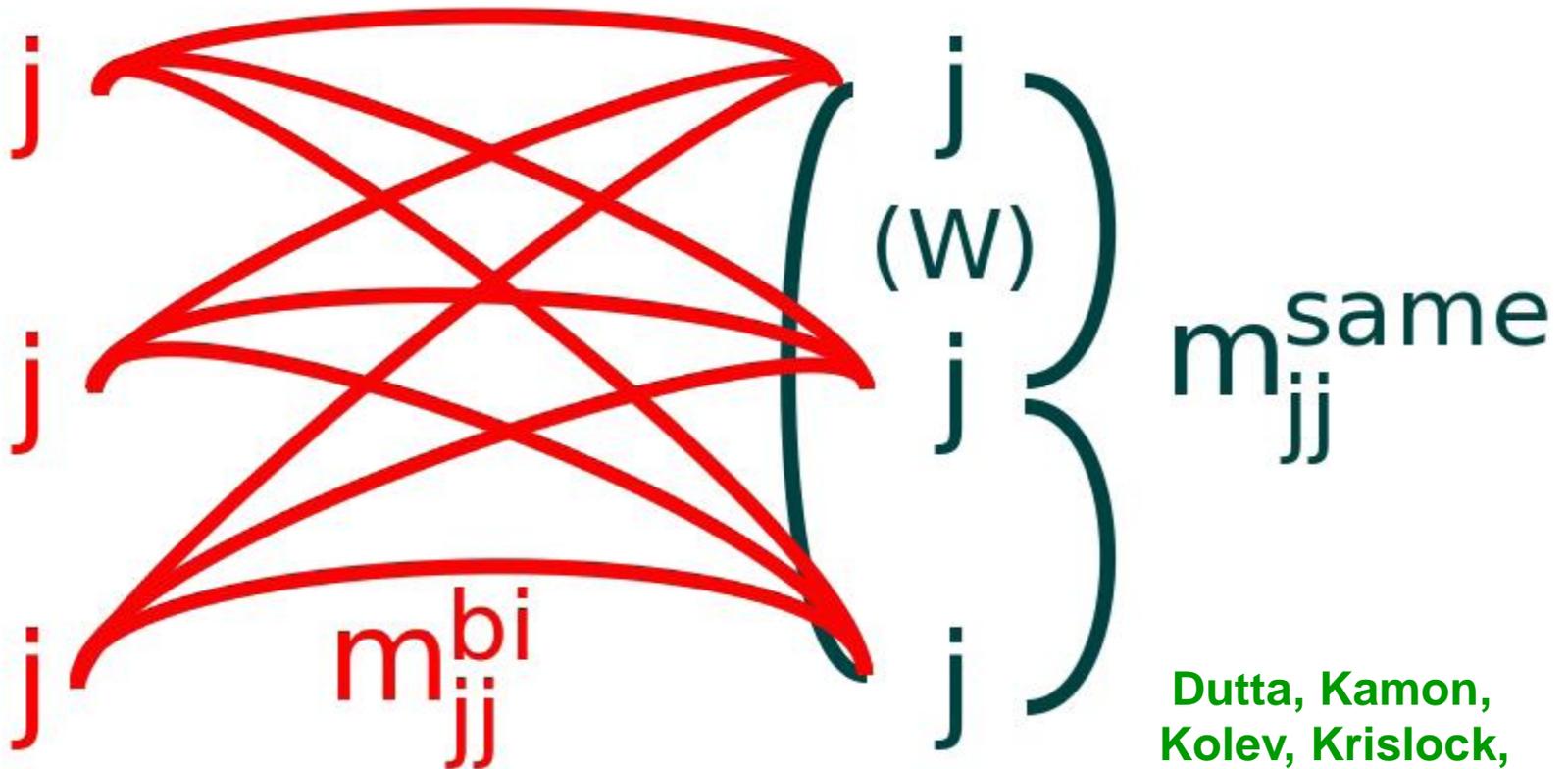
Bi Event Subtraction technique: **BEST**



BEST

Event #n-1

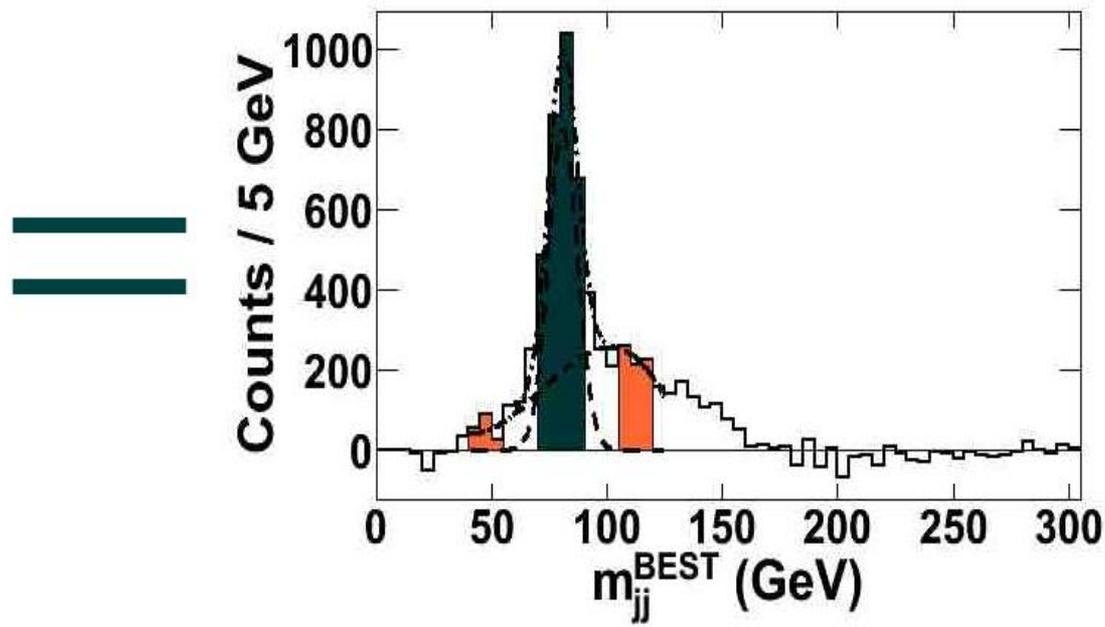
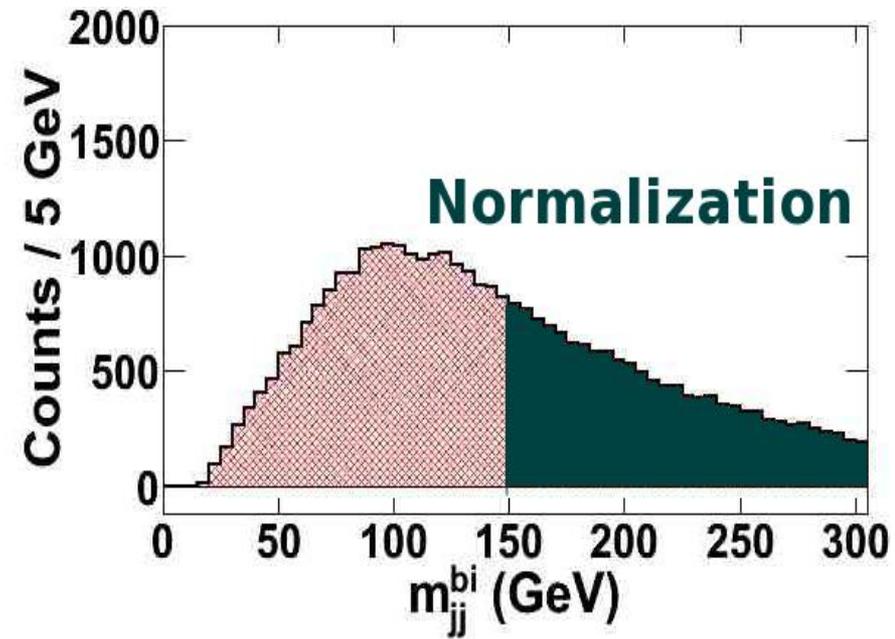
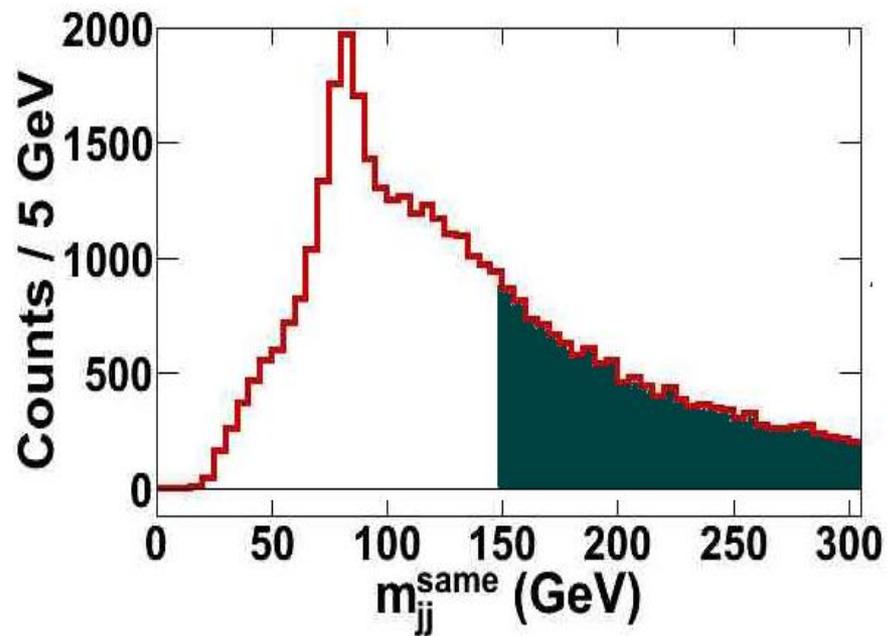
Event #n



Dutta, Kamon,
Kolev, Krislock,

Phys.Lett. B703 (2011) 475

What BEST Looks Like...

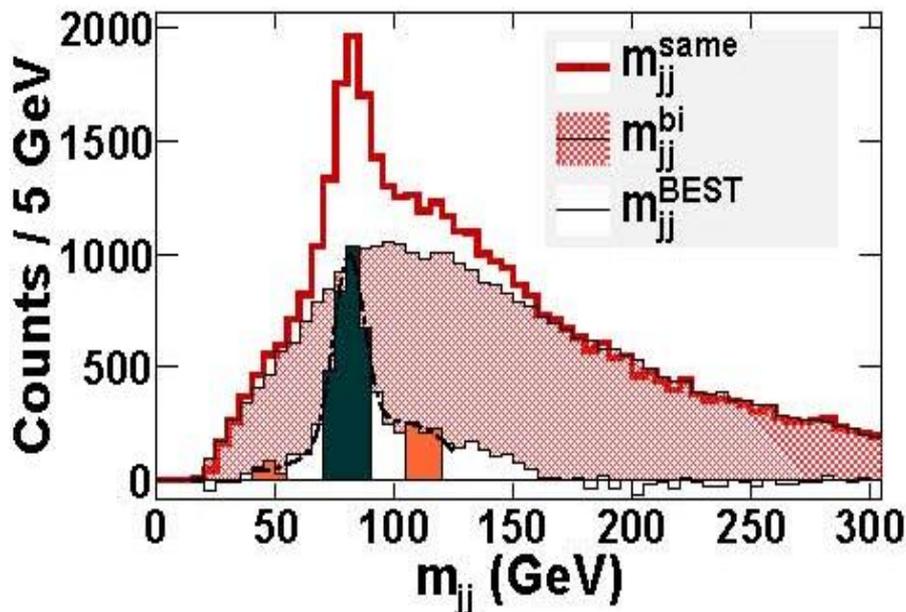


Top reconstruction : BEST

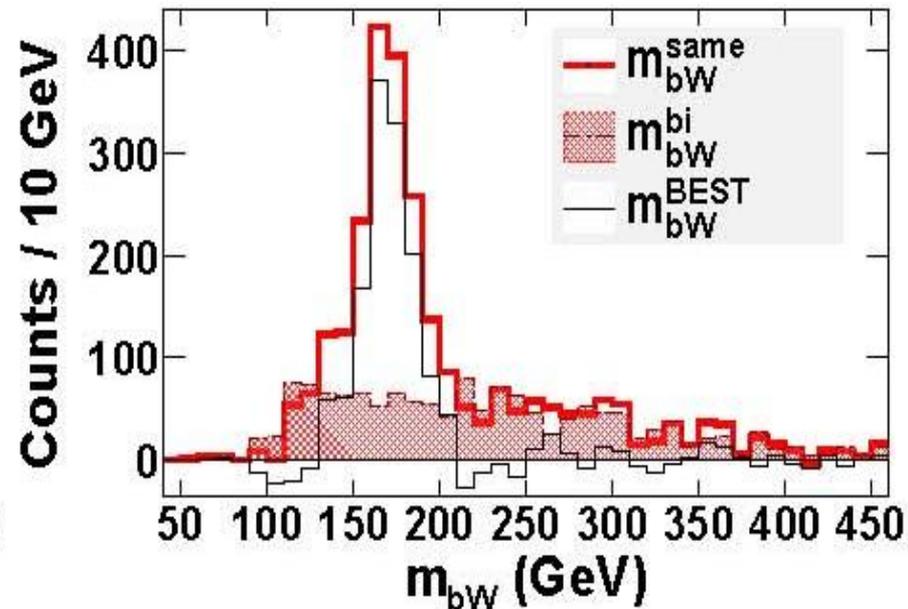
Even with backgrounds, BEST triumphs.

- 7 TeV collision energy @ LHC, 2 fb^{-1} .
- ALPGEN - $t\bar{t}$ signal and W +jets background
- PYTHIA - shower
- PGS - detector

- (i) Number of leptons = 1, where $p_T^l \geq 20 \text{ GeV}$
- (i) Miss. transverse energy > 20 GeV
- (ii) Number of jets, $N \geq 3$, where $p_T^j \geq 30 \text{ GeV}$ and at least one jet has been tightly b-tagged
- (iv) Number of taus, $N_\tau = 0$ for taus with $p_T^\tau \geq 20 \text{ GeV}$

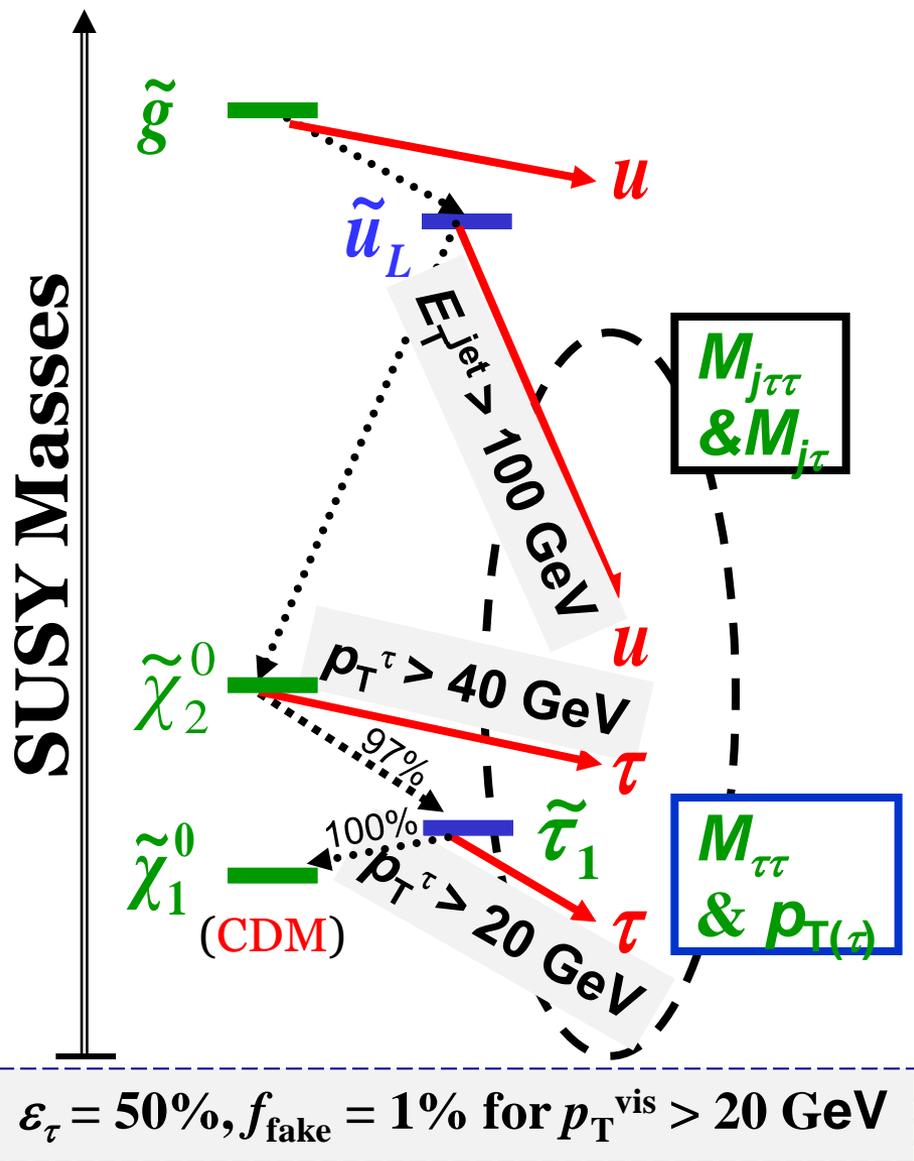


$$m_W = 81.11 \pm 0.32 \text{ GeV}$$



$$m_t = 170.5 \pm 1.5 \text{ GeV}$$

Coann. Region: Final States



Excesses in 3 Final States:

a) $E_T^{\text{miss}} + 4j$

b) $E_T^{\text{miss}} + 2j + 2\tau$

c) $E_T^{\text{miss}} + b + 3j$

Kinematical variables

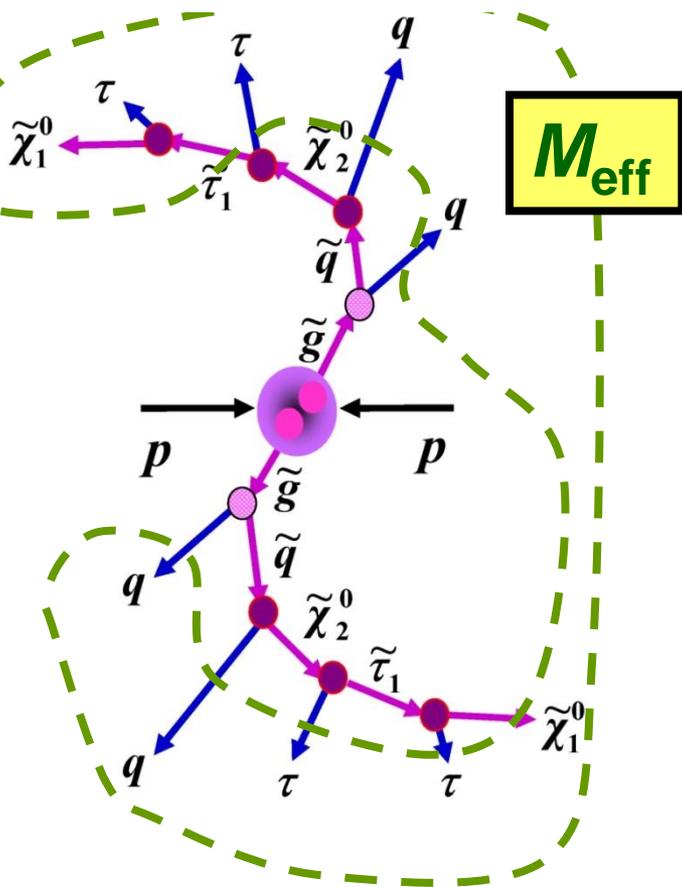
$m_{1/2}, m_0, \tan \beta, A_0$

Different types of final states are needed for different models

a) $E_T^{\text{miss}} + 4j$

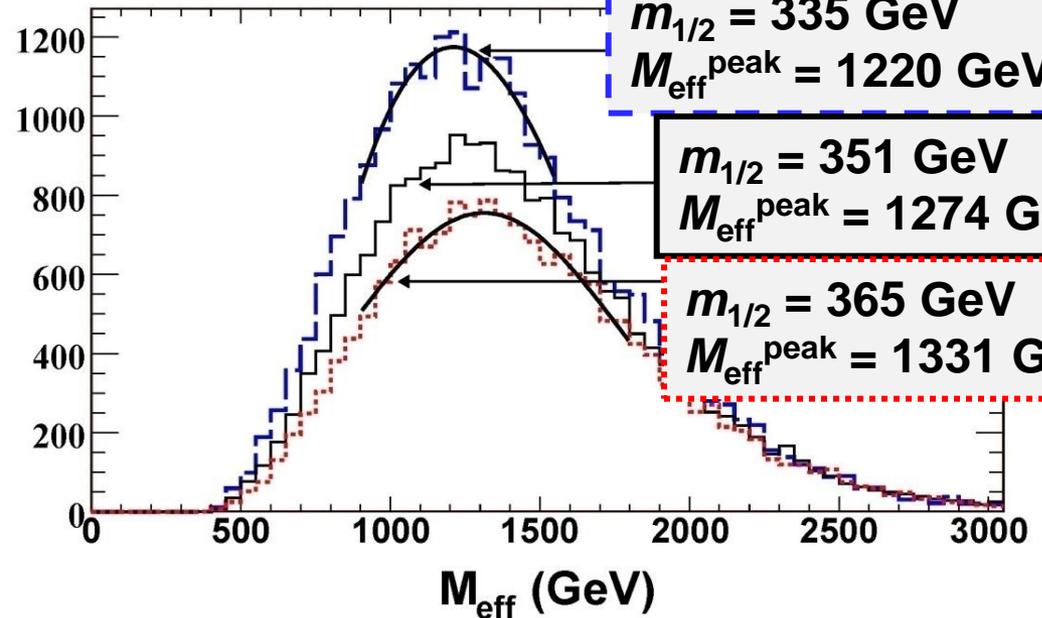
$$M_{\text{eff}} \equiv E_T^{j1} + E_T^{j2} + E_T^{j3} + E_T^{j4} + E_T^{\text{miss}} \quad [\text{No } b \text{ jets; } \varepsilon_b \sim 50\%]$$

e.g., $pp \rightarrow \tilde{g}\tilde{g}$



- $E_T^{j1} > 100$, $E_T^{j2,3,4} > 50$
- No e's, μ 's with $p_T > 20$ GeV
- $M_{\text{eff}} > 400$ GeV;
- $E_T^{\text{miss}} > \max [100, 0.2 M_{\text{eff}}]$

Events / (10 fb⁻¹ X 50 GeV)



c) $E_T^{\text{miss}} + b + 3j$

$$M_{\text{eff}}^{(b)} \equiv E_T^{j1=b} + E_T^{j2} + E_T^{j3} + E_T^{j4} + E_T^{\text{miss}} \quad [j1 = b \text{ jet}]$$

$$E_T^{j1} > 100 \text{ GeV}, \quad E_T^{j2,3,4} > 50 \text{ GeV} \quad [\text{No } e\text{'s}, \mu\text{'s with } p_T > 20 \text{ GeV}]$$

$$M_{\text{eff}}^{(b)} > 400 \text{ GeV}; \quad E_T^{\text{miss}} > \max [100, 0.2 M_{\text{eff}}]$$

$\tan\beta = 48$

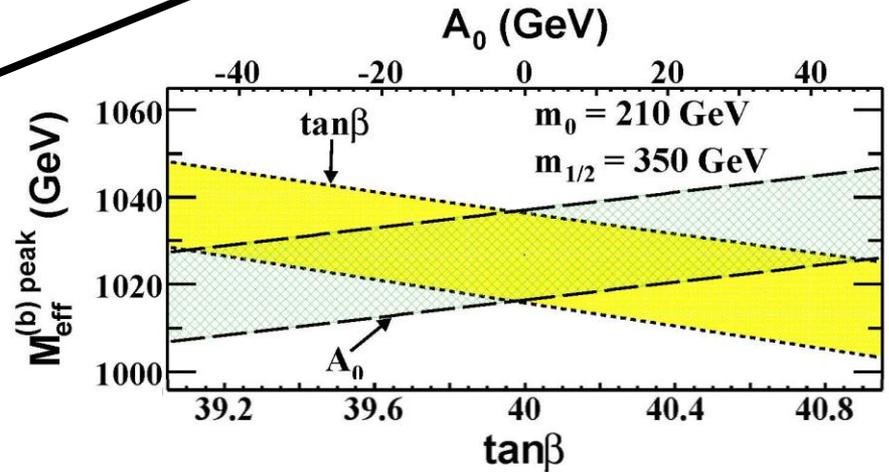
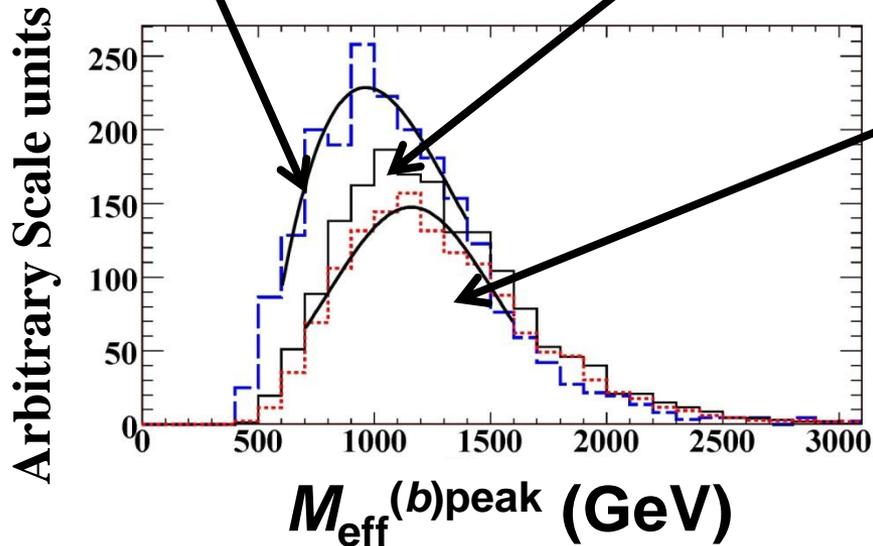
$M_{\text{eff}}^{(b)\text{peak}} = 933 \text{ GeV}$

$\tan\beta = 40$

$M_{\text{eff}}^{(b)\text{peak}} = 1026 \text{ GeV}$

$\tan\beta = 32$

$M_{\text{eff}}^{(b)\text{peak}} = 1122 \text{ GeV}$



$M_{\text{eff}}^{(b)}$ can be used to probe A_0 and $\tan\beta$ without measuring stop and sbottom masses \rightarrow 3rd Gen. squarks gets involved

DM Relic Density in mSUGRA

$$\begin{aligned}
 M_{\tilde{g}} &= 831 \text{ GeV} \\
 M_{\tilde{\chi}_2^0} &= 260 \text{ GeV} \\
 M_{\tilde{\tau}} &= 151.3 \text{ GeV} \\
 M_{\tilde{\chi}_1^0} &= 140.7 \text{ GeV}
 \end{aligned}$$



$$\begin{aligned}
 m_0 &= \\
 m_{1/2} &= \\
 \tan\beta &= \\
 A_0 &= \\
 \text{sgn}(\mu) &> 0
 \end{aligned}$$



$$\Omega_{\tilde{\chi}_1^0} h^2 = Z(m_0, m_{1/2}, \tan\beta, A_0)$$

[1] Established the CA region by detecting low energy τ 's ($p_T^{\text{vis}} > 20 \text{ GeV}$)

[2] Measured 5 SUSY masses ($\Delta M, \tilde{\chi}_1^0, \tilde{\chi}_2^0, \tilde{q}, \tilde{g}$) from

$$M_{j\tau}^{\text{endpoint}} = X_1(m_{1/2}, m_0)$$

$$M_{\tau\tau}^{\text{endpoint}} = X_2(m_{1/2}, m_0, \tan\beta, A_0)$$

$$M_{\text{eff}}^{\text{peak}} = X_3(m_{1/2}, m_0)$$

$$M_{\text{eff,b}}^{\text{peak}} = X_4(m_{1/2}, m_0, \tan\beta, A_0)$$

[3] Determine the dark matter relic density by determining $m_0, m_{1/2}, \tan\beta,$ and A_0

Determining mSUGRA Parameters

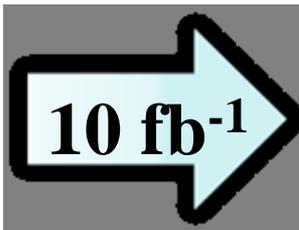
✓ Solved by inverting the following functions:

$$M_{j\tau\tau}^{\text{peak}} = X_1(m_{1/2}, m_0)$$

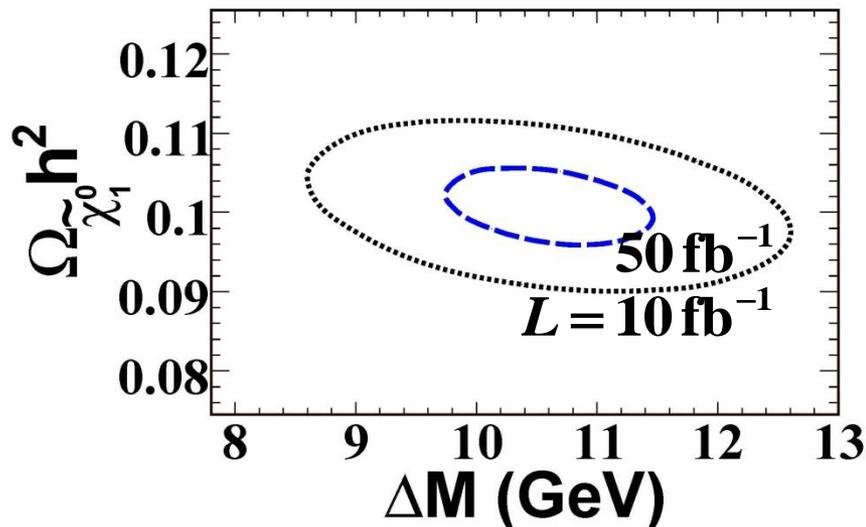
$$M_{\tau\tau}^{\text{peak}} = X_2(m_{1/2}, m_0, \tan\beta, A_0)$$

$$M_{\text{eff}}^{\text{peak}} = X_3(m_{1/2}, m_0)$$

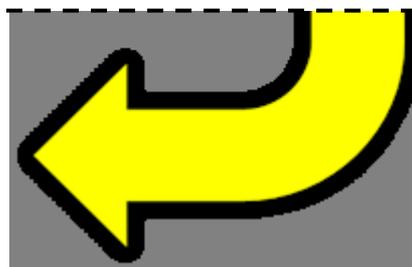
$$M_{\text{eff}}^{(b)\text{peak}} = X_4(m_{1/2}, m_0, \tan\beta, A_0)$$



$$\left\{ \begin{array}{l} m_0 = 210 \pm 5 \\ m_{1/2} = 350 \pm 4 \\ A_0 = 0 \pm 16 \\ \tan\beta = 40 \pm 1 \end{array} \right.$$



$$\Omega_{\tilde{\chi}_1^0} h^2 = Z(m_0, m_{1/2}, \tan\beta, A_0)$$



PRL 100,
(2008),
231802

$$\begin{aligned} \frac{\delta\Omega_{\tilde{\chi}_1^0} h^2}{\Omega_{\tilde{\chi}_1^0} h^2} &= 6.2\% (30 \text{ fb}^{-1}) \\ &= 4.1\% (70 \text{ fb}^{-1}) \end{aligned}$$

$$\frac{\delta\sigma_{\tilde{\chi}_1^0-p}}{\sigma_{\tilde{\chi}_1^0-p}} \approx 7\% (30 \text{ fb}^{-1})$$

Comparison

**ILC analysis:
500 GeV**

LHC

$$m_0 = 210$$

$$\Delta M = 9.5_{-1.0}^{+1.1} \text{ (500 fb}^{-1}\text{)}$$

We need 50fb⁻¹

$$m_{1/2} = 350$$

**Arnowitt, Dutta,
Kamon; PLB 05**



$$A_0 = 0$$



$$\tan \beta = 40$$

**This result was used
in Baltz, Battaglia,
Peskin, Wizansky' 05
to extract relic
density by using
ILC and LHC
(LCC3 point)'05**

**We can determine
 ΔM at the LHC,**

**Arnowitt, Dutta,
Kamon et al,
PRL 08**

Case 2 : Non-U SUGRA

Nature may not be so kind ... Our studies have been done based on a minimal scenario(= mSUGRA)...

Let's consider a non-universal scenario: Higgs non-universality:

m_{Hu}, m_{Hd} m_0 (most plausible extension)

→ easy to explain the DM content:

1) Reduce μ or 2) heavy Higgs/pseudoscalar (A) resonance

Case 1 steps:

1) Reduce Higgs coupling parameter, μ , by increasing m_{Hu} , ...

→ More annihilation (less abundance) → correct values of Ωh^2

2) Find smoking gun signals → Technique to calculate Ωh^2

$$m_{Hu}^2 = m_0^2 (1 + \delta_u^2), m_{Hd}^2 = m_0^2 (1 + \delta_d^2),$$

$$\mu^2 = \left[\frac{\delta_d^2}{\tan^2 \beta} - \delta_u^2 \frac{(1 + D_0)}{2} \right] m_0^2 + \dots$$

Where $D_0 < 0.23$

For low and intermediate $\tan\beta$...

Reference Point

Parameters at the GUT scale:

- $m_0 = 360 \text{ GeV}$, $m_{1/2} = 500 \text{ GeV}$, $A_0 = 0 \text{ GeV}$, $\tan \beta = 40$
- Non-universal Higgs: $m_{H_u} = 732 \text{ GeV}$, $m_{H_d} = 732 \text{ GeV}^*$

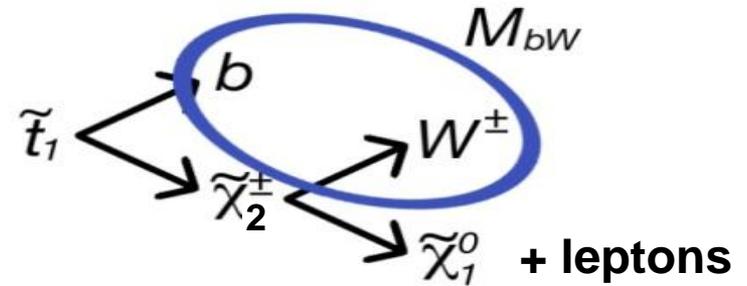
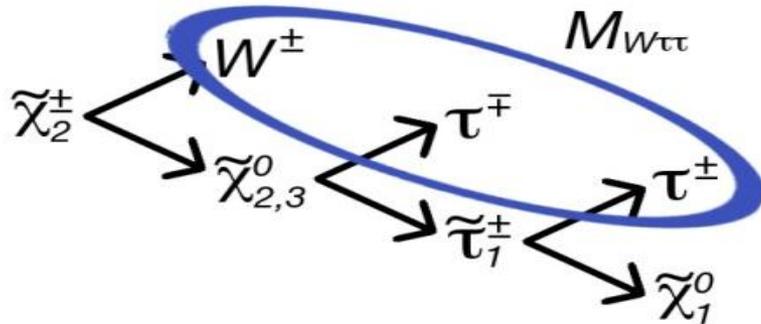
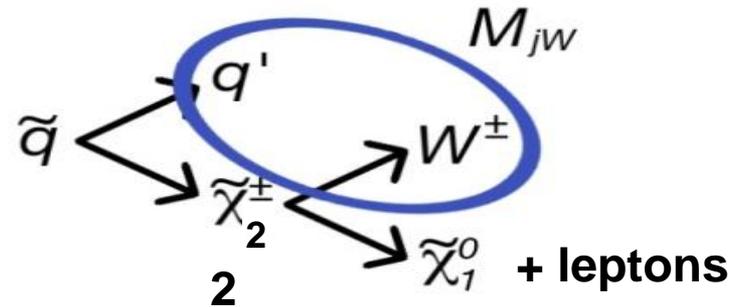
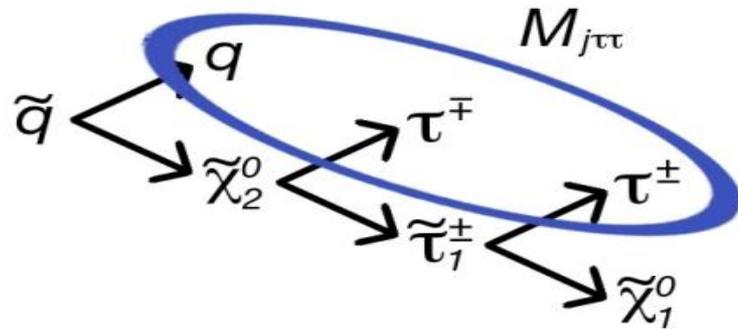
$$\Omega h^2 = 0.112$$

SUSY masses (in GeV):

\tilde{g}	\tilde{U}_L \tilde{U}_R	\tilde{t}_2 \tilde{t}_1	\tilde{b}_2 \tilde{b}_1	\tilde{e}_L \tilde{e}_R	$\tilde{\tau}_2$ $\tilde{\tau}_1$	$\tilde{\chi}_4^0$ $\tilde{\chi}_3^0$ $\tilde{\chi}_2^0$ $\tilde{\chi}_1^0$	$\tilde{\chi}_2^\pm$ $\tilde{\chi}_1^\pm$
						432	
1161	1114 1076	992 780	989 946	494 407	446 255	317 293 199	428 292

Decays at Reference Point

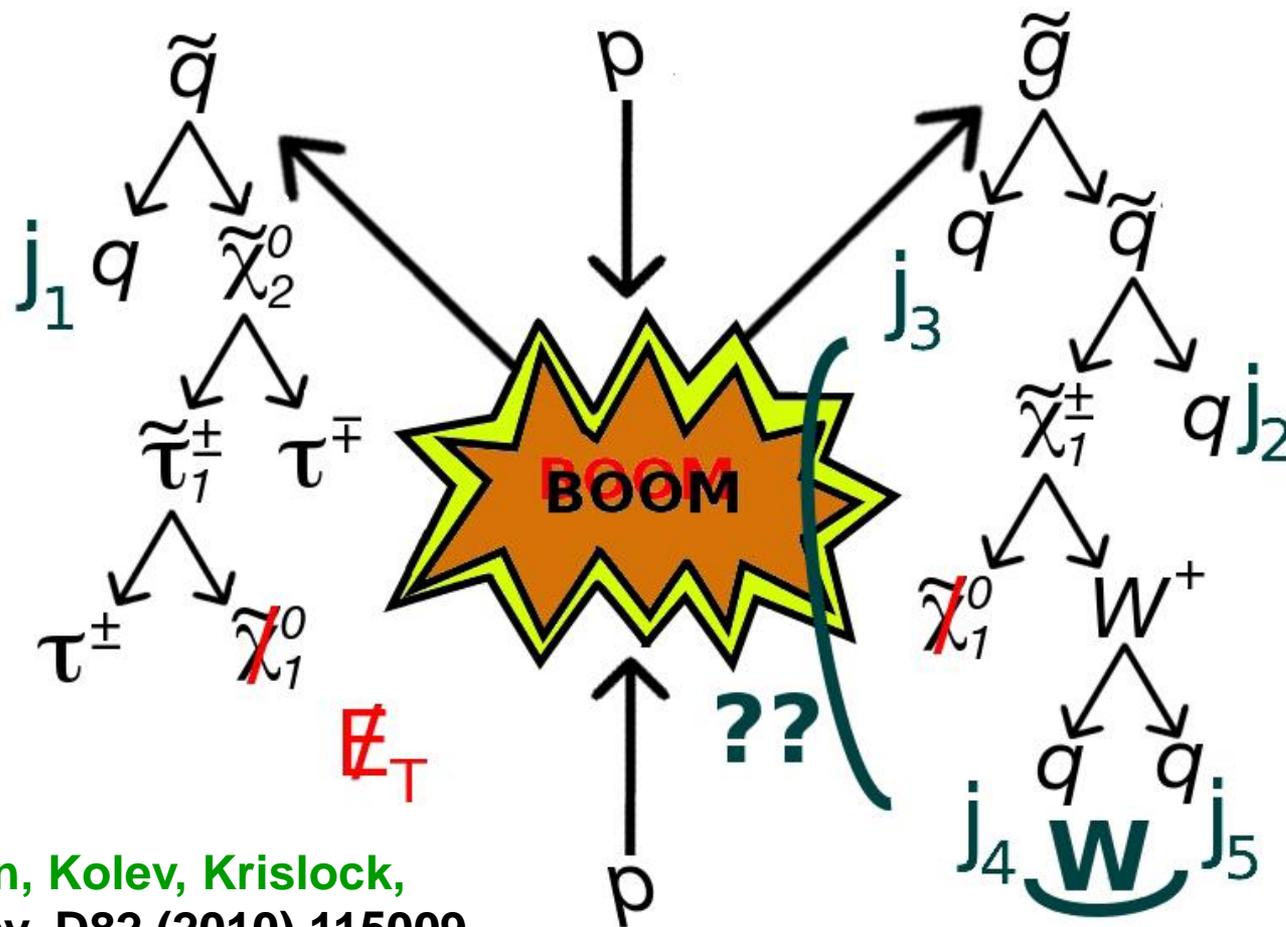
Benchmark Point: Characteristic Decays



In the non-universal scenario: We use W + jets etc.

BEST and nuSUGRA...

In this scenario we have W's in the final states:



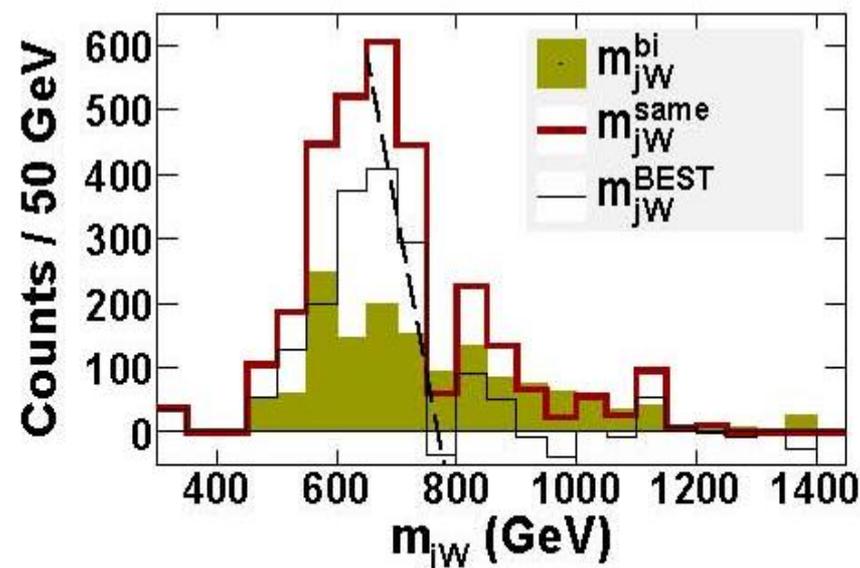
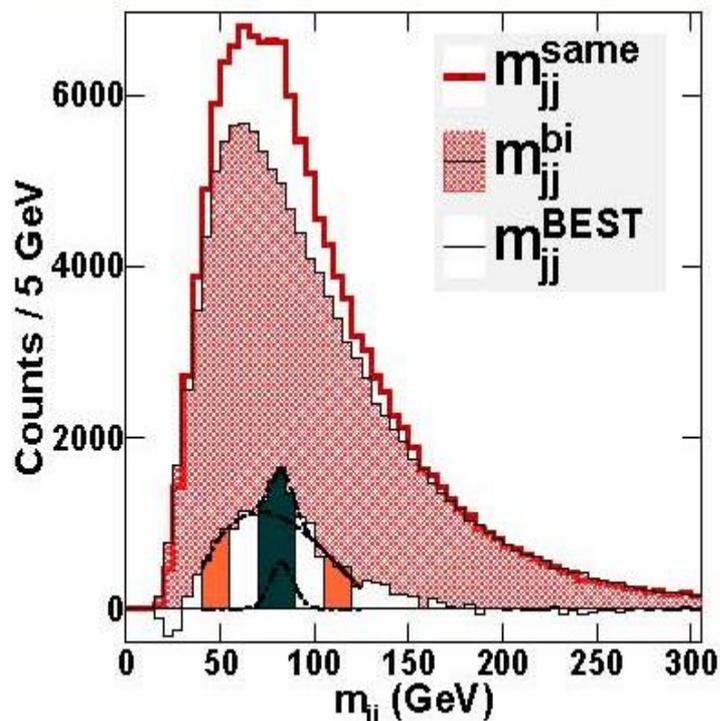
Dutta, Kamon, Kolev, Krislock,
Oh, Phys.Rev. D82 (2010) 115009

End Point Techniques with BEST

Even with backgrounds on top of SUSY, BEST triumphs.

- 14 TeV collision energy @ LHC, 100 fb^{-1} .
- nuSUGRA: $m_0 = 360 \text{ GeV}$, $m_{1/2} = 500 \text{ GeV}$, $\tan \beta = 40$, $A_0 = 0$, and $m_H = 732 \text{ GeV}$.
- SM: $t\bar{t}$, W +Jets, and Z +Jets.

- $N_{\text{jet}} > 4$, $p_T > 30$
- $E_T^{j1,2} > 100$, $E_T^{\text{miss}} > 180$
- $E_T^{\text{miss}} + E_T^{j1} + E_T^{j2} > 600$
- No e 's, μ 's with $p_T > 5$

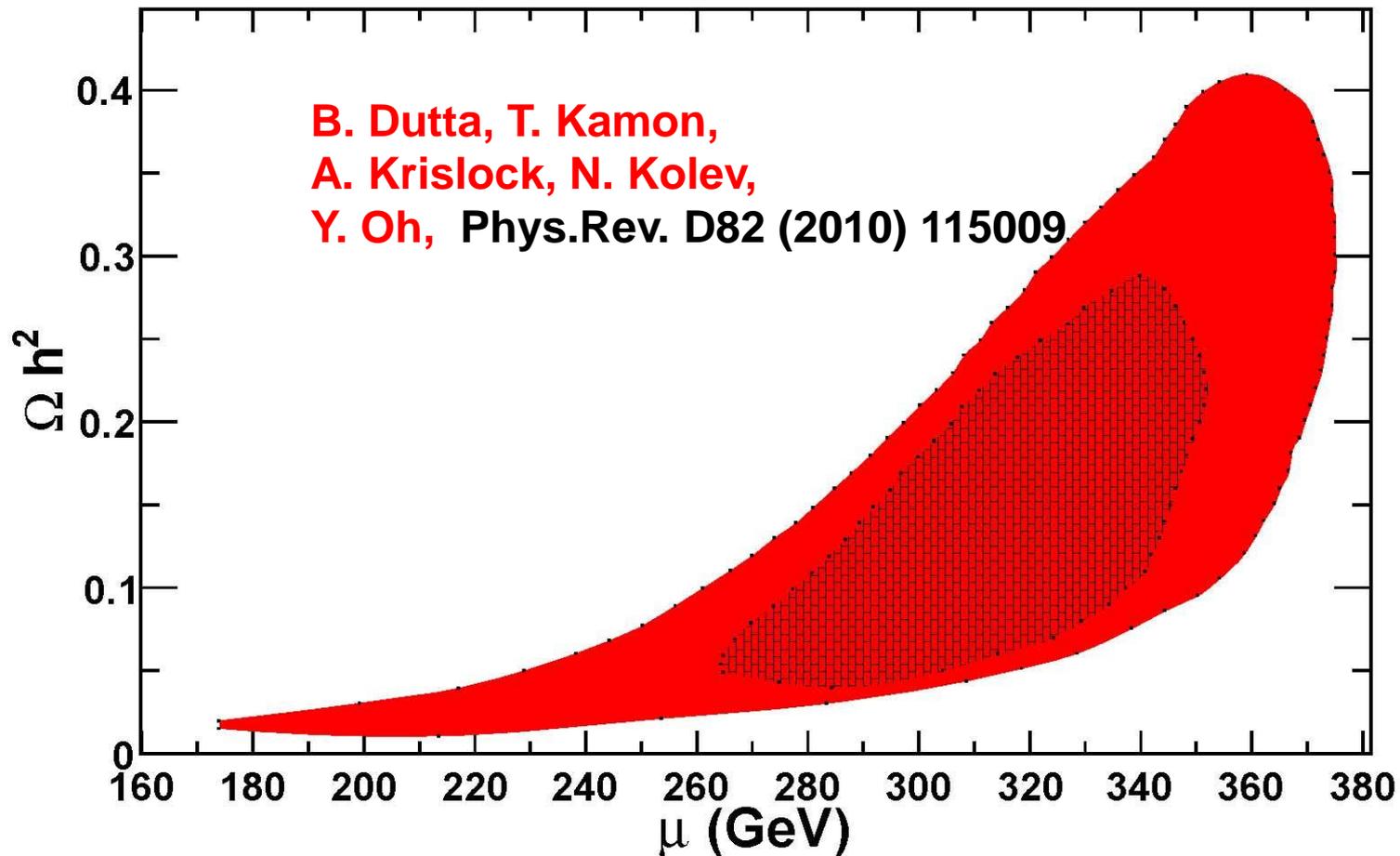


$$m_{jW}^{\text{max}} = 769 \pm 18 \text{ GeV}$$

Significance improves 5 times with BEST

Relic Density

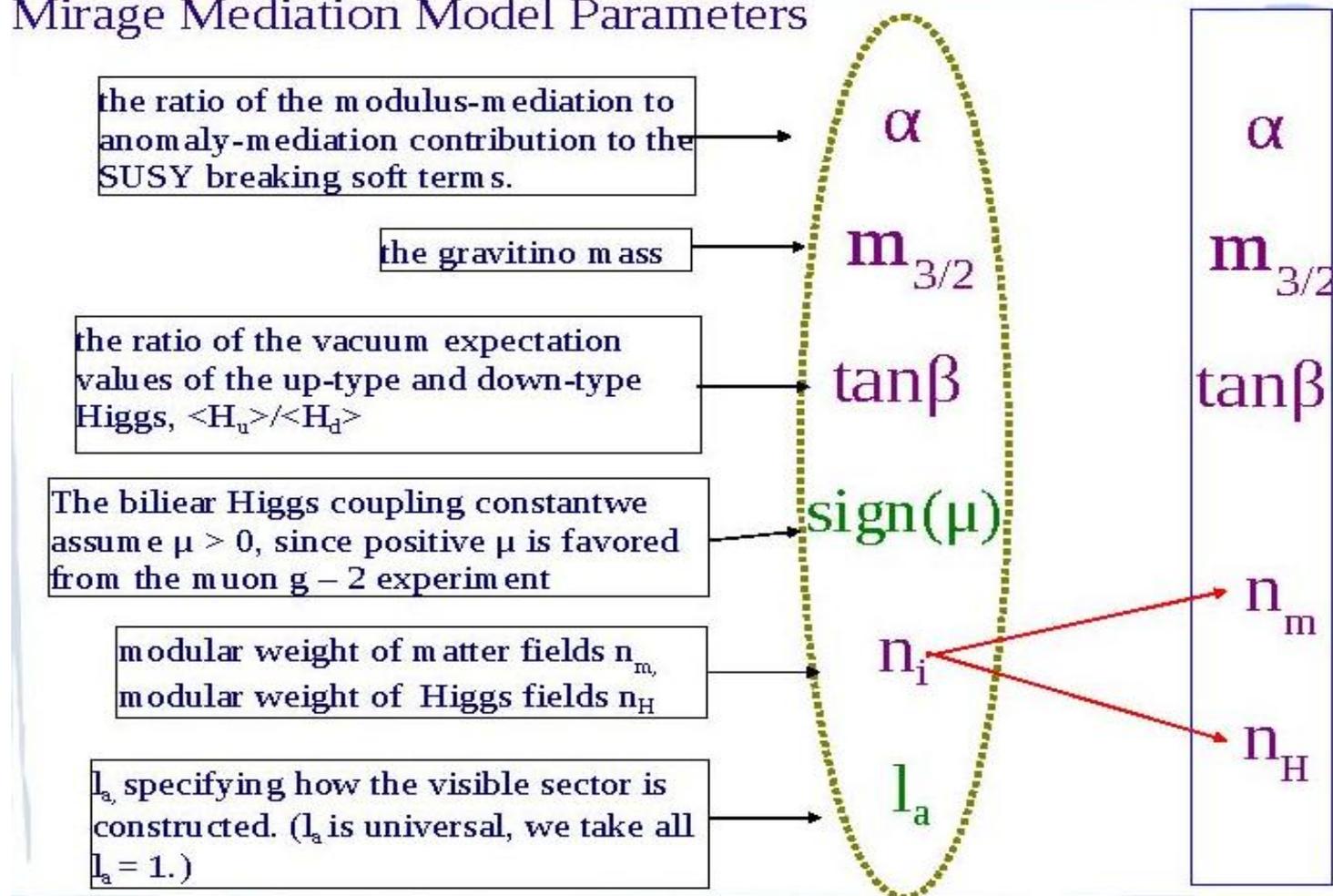
\mathcal{L} (fb $^{-1}$)	$m_{1/2}$ (GeV)	m_H (GeV)	m_0 (GeV)	A_0 (GeV)	$\tan\beta$	μ (GeV)	$\Omega_{\tilde{\chi}_1^0} h^2$
1000	500 ± 3	727 ± 10	366 ± 26	3 ± 34	39.5 ± 3.8	321 ± 25	$0.094^{+0.107}_{-0.038}$
100	500 ± 9	727 ± 13	367 ± 57	0 ± 73	39.5 ± 4.6	331 ± 48	$0.088^{+0.168}_{-0.072}$
Syst.	± 10	± 15	± 56	± 66	± 4.5	± 48	$^{+0.175}_{-0.072}$



Case 3 : Mirage Mediation

Soft masses: Moduli mediation + anomaly mediation

Mirage Mediation Model Parameters



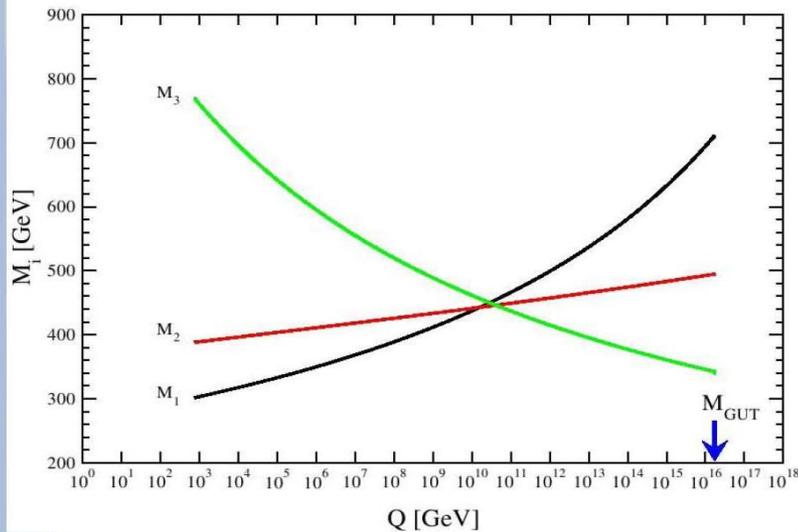
Mirage Mediation

Mirage Unification of the Gaugino Masses

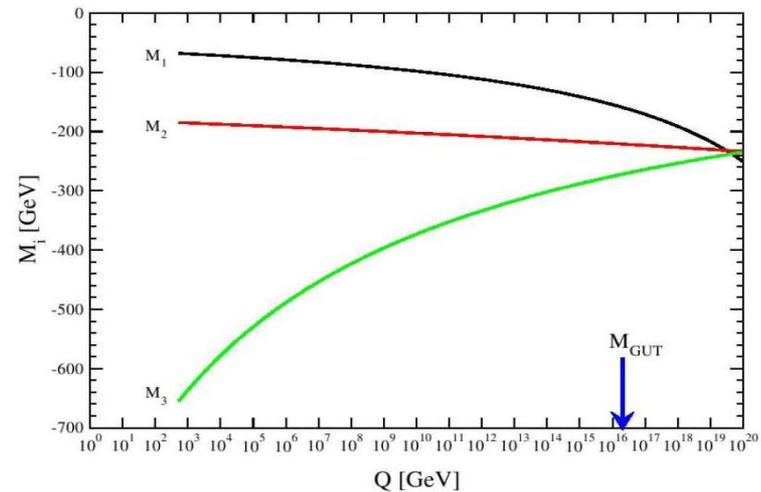
Gaugino masses are unified at the mirage unification scale.

The mirage unification scale is given as: $\mu_{\text{mir}} = M_{\text{GUT}} e^{-8 \pi^2 / \alpha}$

$n_H=1, n_m=1/2 : \alpha=6, m_{3/2}=12 \text{ TeV}, \tan\beta=10, \mu>0, m_t=175 \text{ GeV}$



$n_H=0, n_m=1 : \alpha=-10, m_{3/2}=4 \text{ TeV}, \tan\beta=10, \mu>0, m_t=175 \text{ GeV}$



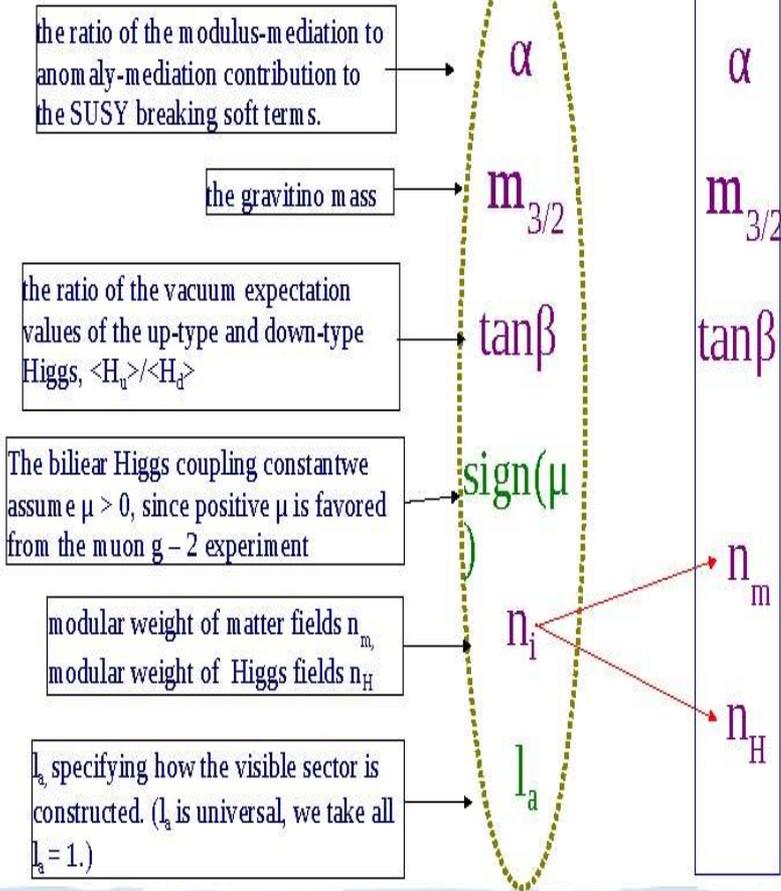
Howard Baer, Eun-Kyung Park, etc., arXiv:hep-ph/0703024v2

$$M_a(\mu) = \frac{m_{3/2}}{16 \pi^2} \alpha \left[1 - \frac{1}{8 \pi^2} b_a g_a^2(\mu) \ln \left(\frac{\mu_{\text{mir}}}{\mu} \right) \right]$$

Non universality of the gaugino masses at the GUT scale

Mirage Mediation

Mirage Mediation Model Parameters



$$m_{\tilde{g}} = \text{function}(\alpha, m_{3/2})$$

$$m_{\tilde{\chi}_1^0} = \text{function}(\alpha, m_{3/2})$$

$$m_{\tilde{q}} = \text{function}(\alpha, m_{3/2}, n_m)$$

$$m_{\tilde{\chi}_2^0} = \text{function}(\alpha, m_{3/2}, n_m, n_H)$$

$$m_{\tilde{\tau}_1} = \text{function}(\alpha, m_{3/2}, n_m, n_H, \tan\beta)$$



$$\log(\text{PT}_{\tau}^{\text{slope}}) = \text{function}(m_{\tilde{\tau}_1}, m_{\tilde{\chi}_1^0})$$

$$M_{\tau\tau}^{\text{end}} = \text{function}(m_{\tilde{\chi}_2^0}^2, m_{\tilde{\tau}_1}^2, m_{\tilde{\chi}_1^0}^2)$$

$$M_{j\tau}^{\text{end}} = \text{function}(m_{\tilde{q}}, m_{\tilde{\chi}_2^0}^2, m_{\tilde{\tau}_1}^2)$$

$$M_{j\tau\tau}^{\text{end}} = \text{function}(m_{\tilde{q}}, m_{\tilde{\chi}_2^0}^2, m_{\tilde{\chi}_1^0}^2)$$

$$M_{\text{eff}}^{\text{peak}} \simeq \text{function}(m_{\tilde{g}}, m_{\tilde{q}})$$

Mirage Mediation

Dark matter allowed regions:

- 1. Stop Coannihilation**
- 2. Stau Coannihilation**
- 3. Higgsino domination**
- 4. Wino domination**
- 5. Pseudo scalar Higgs resonance**

Two main goals: Gaugino masses, DM content

Mirage Mediation

Typical stau-neutralino coannihilation region

Parameter	Value	Particle	Mass	Particle	Mass	Particle	Mass
α	7.5	\tilde{d}_L	845.49	\tilde{e}_L	426.91	$\tilde{\chi}_1^0$	284.17
$m_{3/2}$	10000	\tilde{d}_R	813.52	\tilde{e}_R	367.70	$\tilde{\chi}_2^0$	389.17
n_m	0.5	\tilde{u}_L	841.39	$\tilde{\tau}_1$	309.75	$\tilde{\chi}_3^0$	548.88
n_H	1.0	\tilde{u}_R	815.27	$\tilde{\tau}_2$	425.68	$\tilde{\chi}_4^0$	569.04
$\tan\beta$	30	\tilde{b}_1	735.87			$\tilde{\chi}_1^\pm$	389.32
		\tilde{b}_2	791.30			$\tilde{\chi}_2^\pm$	568.10
		\tilde{t}_1	600.23			\tilde{g}	897.55
		\tilde{t}_2	810.20				

Observable	Value	100 fb ⁻¹ Stat.
$M_{\tau\tau}^{\text{end}}$	90.70	± 0.54
$M_{j\tau\tau}^{\text{end}}$	479.53	± 3.45
$\text{slope}(p_{T,\tau})$	-0.0849	± 0.0041
$M_{\text{eff}}^{\text{peak}}$	1257.26	± 10.33
$M_{j\tau}^{\text{end}}$	448.40	± 16.20

Mirage Mediation

Particle	Mass	100 fb^{-1} Stat.
\tilde{g}	895	$-35, +50$
\tilde{q}_L	845	$-36, +24$
$\tilde{\chi}_2^0$	388	$-9, +25$
$\tilde{\tau}$	298	$-8, +8$
$\tilde{\chi}_1^0$	274	$-10, +10$

Parameter	Value	Stat.
α	7.42	± 0.58
$m_{3/2}$	10171	± 882
n_m	0.52	± 0.09
n_H	1.17	$-0.07, +0.22$
$\tan\beta$	33.1	± 7.8

$$\Omega h^2 = 0.17_{-0.13}^{+0.12} .$$

Mirage Mediation

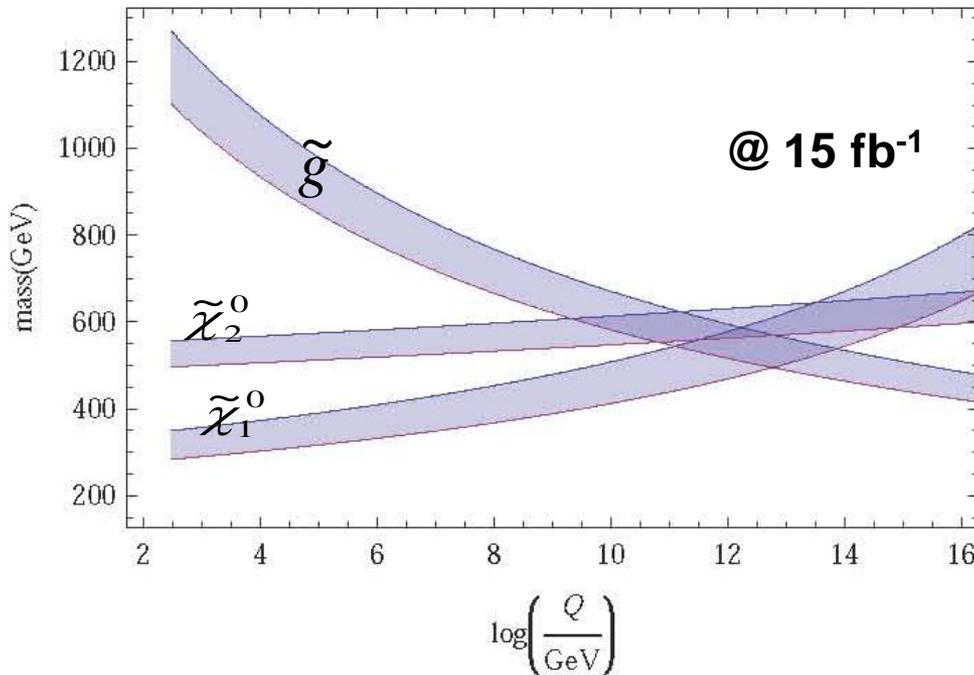
Glينو mass ~ 1.1 TeV

Parameter	Value	Particle	Mass
α	10	\tilde{g}	1183
$m_{3/2}$	9500	$\tilde{\chi}_2^0$	499
n_m	0.5	$\tilde{\chi}_1^0$	337
n_H	1.0	$\tilde{\tau}$	361
$\tan\beta$	37	\tilde{q}	1119

$P_{T_low}(m_{\tau 1}, m_{\chi 2}, m_{\chi 1});$
 $m_{\tau\tau}(m_{\tau 1}, m_{\chi 2}, m_{\chi 1});$
 $M_{eff}(m_{gluino}, m_{\chi 1})$

Particle	Mass	15fb^{-1} Stat.
\tilde{g}	1186	± 84
$\tilde{\chi}_2^0$	527	$-30, +60$
$\tilde{\chi}_1^0$	317	± 33

Input from the experimental measurements



Values of the masses at the GUT scale

Using all the observables for this point: $\Omega h^2 = 0.05^{+0.21}_{-0.04}$.

Mirage Mediation

Typical stop-neutralino coannihilation region

$m_{3/2}$	α	$\tan\beta$	n_m	n_H
14000	4.5	30	0	0.5

Mass Spectrum
(GeV)

\tilde{g}	\tilde{u}_L	\tilde{b}_2	\tilde{t}_2	\tilde{e}_L	$\tilde{\tau}_2$	$\tilde{\chi}_2^0$
	\tilde{u}_R	\tilde{b}_1	\tilde{t}_1	\tilde{e}_R	$\tilde{\tau}_1$	$\tilde{\chi}_1^0$
650	648	596	617	437	418	338
	635	521	336	411	315	286

Mirage Mediation

Particle	Mass	50 fb ⁻¹ Stat.	100 fb ⁻¹ Stat.
\tilde{g}	646	-14,+19	-11,+14
\tilde{q}_L	638	-34,+42	-23,+39
$\tilde{\tau}$	318	-3,+3	-3,+3
$\tilde{\chi}_2^0$	333	-7,+11	-6,+8
$\tilde{\chi}_1^0$	276	-8,+13	-7,+10

Particle	Mass	50 fb ⁻¹ Stat.	100 fb ⁻¹ Stat.
\tilde{b}	531	-60,+60	-47,+47
\tilde{t}	326	-5,+8	-4,+7

Parameter	Value	50 fb ⁻¹ Stat.	100 fb ⁻¹ Stat.
α	4.58	± 0.21	± 0.14
$m_{3/2}$	13717	± 688	± 517
n_m	0.106	± 0.015	± 0.015
n_H	0.578	± 0.095	± 0.091
$\tan\beta$	28.76	± 1.65	± 1.36

**Dutta, Kamon,
Krislock, Sinha, Wang,
Phys.Rev. D85 (2012) 115007**

$$\Omega h^2 = 0.096 \pm 0.029$$

Mirage Mediation

Parameter	Value	Particle	Mass
α	3.8	\tilde{g}	1187
$m_{3/2}$	34800	$\tilde{\chi}_2^0$	740
n_m	0.0	$\tilde{\chi}_1^0$	666
n_H	0.5	$\tilde{\tau}$	721
$\tan\beta$	28	\tilde{q}	1189

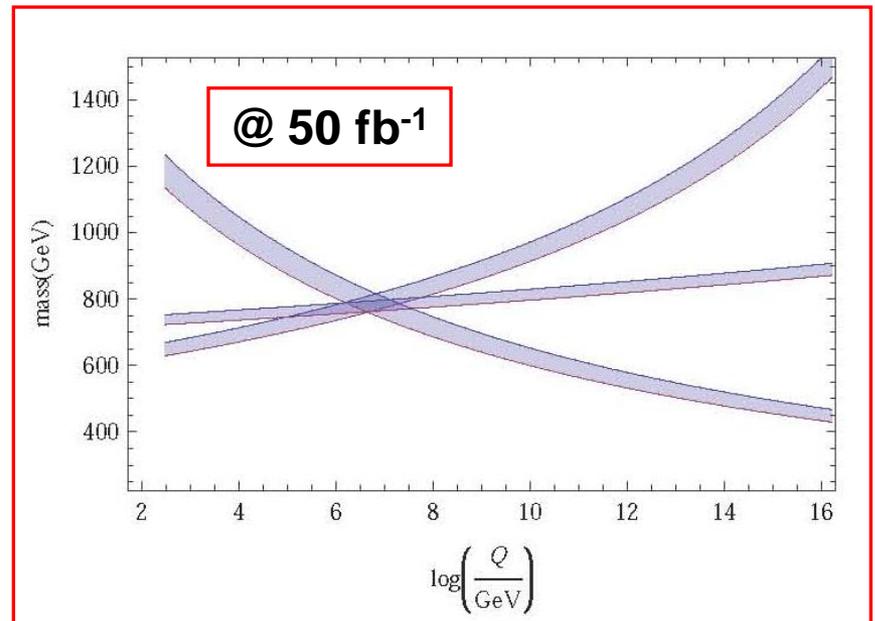
Particle	Mass	Stat.	
\tilde{t}	690	± 6	@ 200 fb⁻¹
\tilde{b}	1002	± 126	
$\tilde{\tau}$	717	± 10	
\tilde{q}	1133	-132, +167	

$$\Omega h^2 = 0.23 \pm 0.13.$$

$\mathbf{P}_{T_sum} (m_{\tau 1}, m_{\chi 2}, m_{\chi 1});$
 $\mathbf{P}_{T_diff} (m_{\tau 1}, m_{\chi 2}, m_{\chi 1});$
 $\mathbf{m}_{\tau\tau} (m_{\tau 1}, m_{\chi 2}, m_{\chi 1});$
 $\mathbf{M}_{eff} (m_{gluino}, m_{\chi 1})$

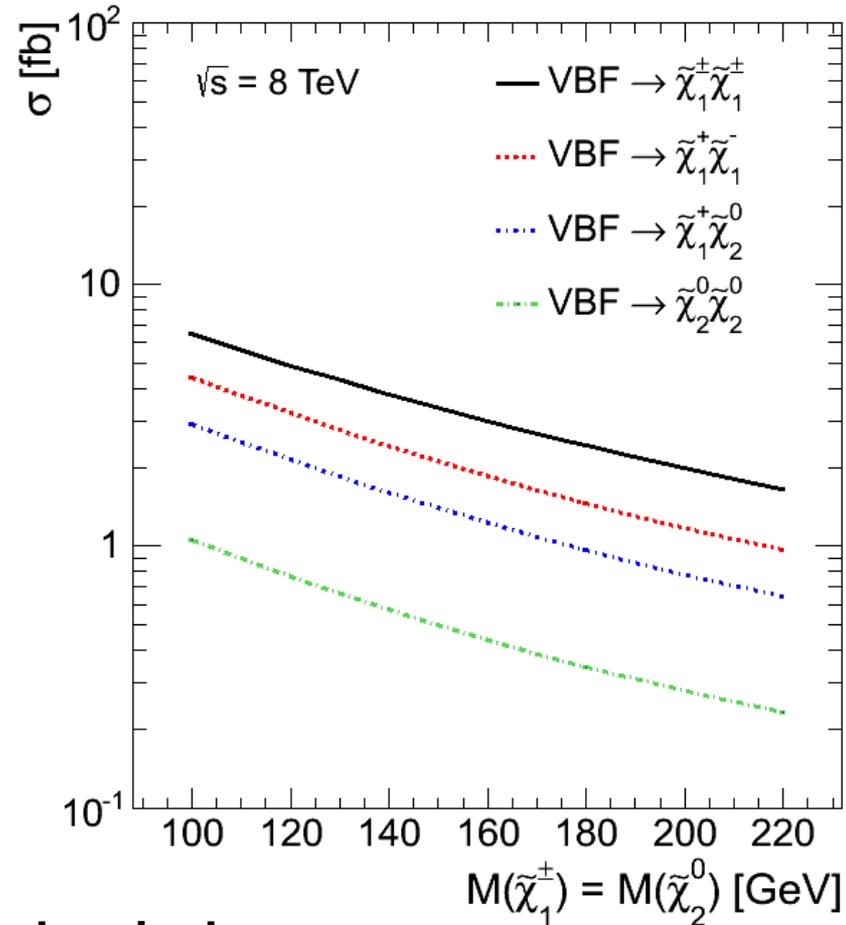
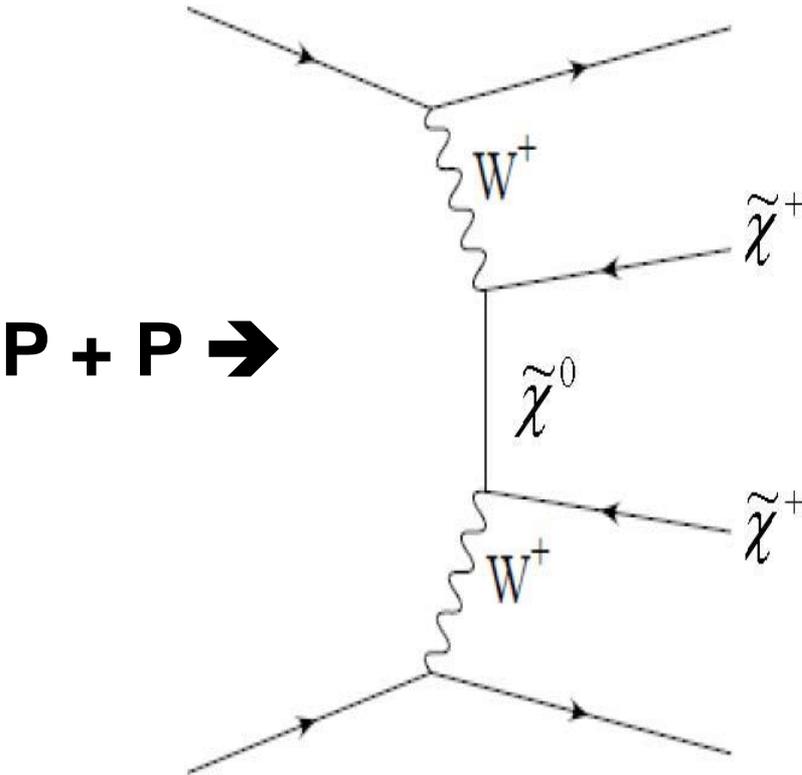
Particle	Mass	50fb ⁻¹	Stat.
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\tilde{g}	1181	± 50
$\tilde{\chi}_2^0$	738	± 15
$\tilde{\chi}_1^0$	649	± 20



2. DM at the LHC Via VBF

Direct probes of charginos, neutralinos and sleptons



Two high E_T forward jets in opposite hemispheres
with large dijet invariant mass

DM at the LHC Via VBF

The decay modes of charginos, neutralinos:

$$\tilde{\chi}_1^\pm \rightarrow \tilde{\tau}_1^\pm \nu \rightarrow \tau^\pm \tilde{\chi}_1^0 \nu$$

Signal: $\geq 2j + 2\tau +$ missing energy

$$\tilde{\chi}_2^0 \rightarrow \tilde{\tau}_1^\pm \tau^\mp \rightarrow \tau^\pm \tau^\mp \tilde{\chi}_1^0$$

Benchmark scenario:

$$m_{\tilde{\chi}_1^\pm} \sim m_{\tilde{\chi}_2^0} = 180 \text{ GeV}, m_{\tilde{\tau}_1} - m_{\tilde{\chi}_1^0} = 30 \text{ GeV}, m_{\tilde{\chi}_1^0} = 90 \text{ GeV}$$

Also:

$$\tilde{\chi}_1^\pm \rightarrow \tilde{l}_1^\pm \nu \rightarrow l^\pm \tilde{\chi}_1^0 \nu$$

$$\tilde{\chi}_2^0 \rightarrow \tilde{l}_1^\pm l^\mp \rightarrow l^\pm l^\mp \tilde{\chi}_1^0$$

Signal: $\geq 2j + 2\mu +$ missing energy

Signal: $\geq 2j+2\tau$ +missing energy

2 jets each with $p_T > 50$ GeV, leading $p_T > 75$ GeV
 $|\Delta\eta(j_1, j_2)| > 4.2$, $\eta_{j_1} \eta_{j_2} < 0$, $M_{j_1 j_2} > 650$ GeV

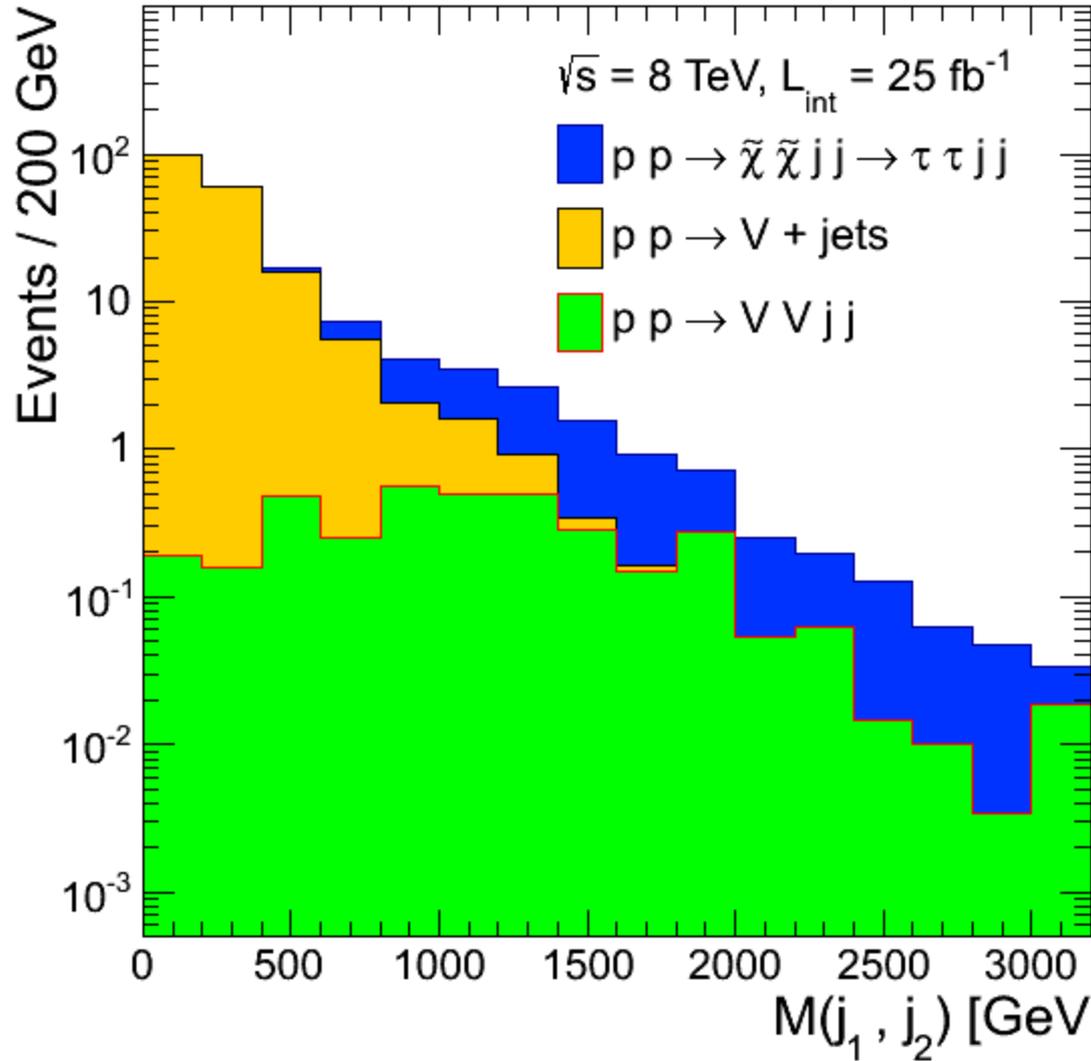
Signal: $\geq 2j + 2\tau +$ missing energy

$m_{\tilde{\chi}_1^\pm} \sim m_{\tilde{\chi}_2^0} = 180 \text{ GeV}$, $\sqrt{s} = 8 \text{ TeV}$, Lum: 25 fb^{-1}

	Signal	Z+jets	W+jets	WW	WZ
VBF cuts	4.61	10.9	3.70×10^3	97.0	19.0
$\cancel{E}_T > 75$	4.33	0.27	5.29×10^2	17.6	3.45
2τ , inclusive	0.45	0.06	0.23	0.09	0.04
(S/\sqrt{B})			3.47		
$\tau^\pm \tau^\pm$	0.21	0	0.11	0.02	0.01
(S/\sqrt{B})			2.91		
$\tau^\pm \tau^\mp$	0.24	0.06	0.12	0.07	0.03
(S/\sqrt{B})			2.27		

Two τ 's with $p_T > 20$ GeV in $\eta < 2.1$, with $\Delta R(\tau\tau) > 0.3$. All τ 's are hadronic.
 The τ ID efficiency is assumed to be 55% and the jet $\rightarrow \tau$ Mis-identification rate is taken to be 1%,

Signal: $\geq 2j+2\tau$ +missing energy



Signal: $\geq 2j+2\mu$ +missing energy

2 jets each with $p_T > 50$ GeV, leading $p_T > 75$ GeV
 $|\Delta\eta(j_1, j_2)| > 4.2$, $\eta_{j_1} \eta_{j_2} < 0$, $M_{j_1 j_2} > 650$ GeV

Signal: $\geq 2j + 2\mu$ + missing energy

$$m_{\tilde{\chi}_1^\pm} \sim m_{\tilde{\chi}_2^0} = 180 \text{ GeV},$$

	Signal	Z+jets	W+jets	WW	WZ
VBF cuts	4.61	10.9	3.70×10^3	0.97×10^2	19.0
$\cancel{E}_T > 75$	4.33	0.27	5.29×10^2	17.6	3.45
2μ , inclusive	1.83	0.15	0	0.12	0.19
(S/\sqrt{B})			13.5		
$\mu^\pm \mu^\pm$	0.87	0	0	0.03	0.05
(S/\sqrt{B})			15.4		
$\mu^\pm \mu^\mp$	0.96	0.15	0	0.09	0.14
(S/\sqrt{B})			7.80		

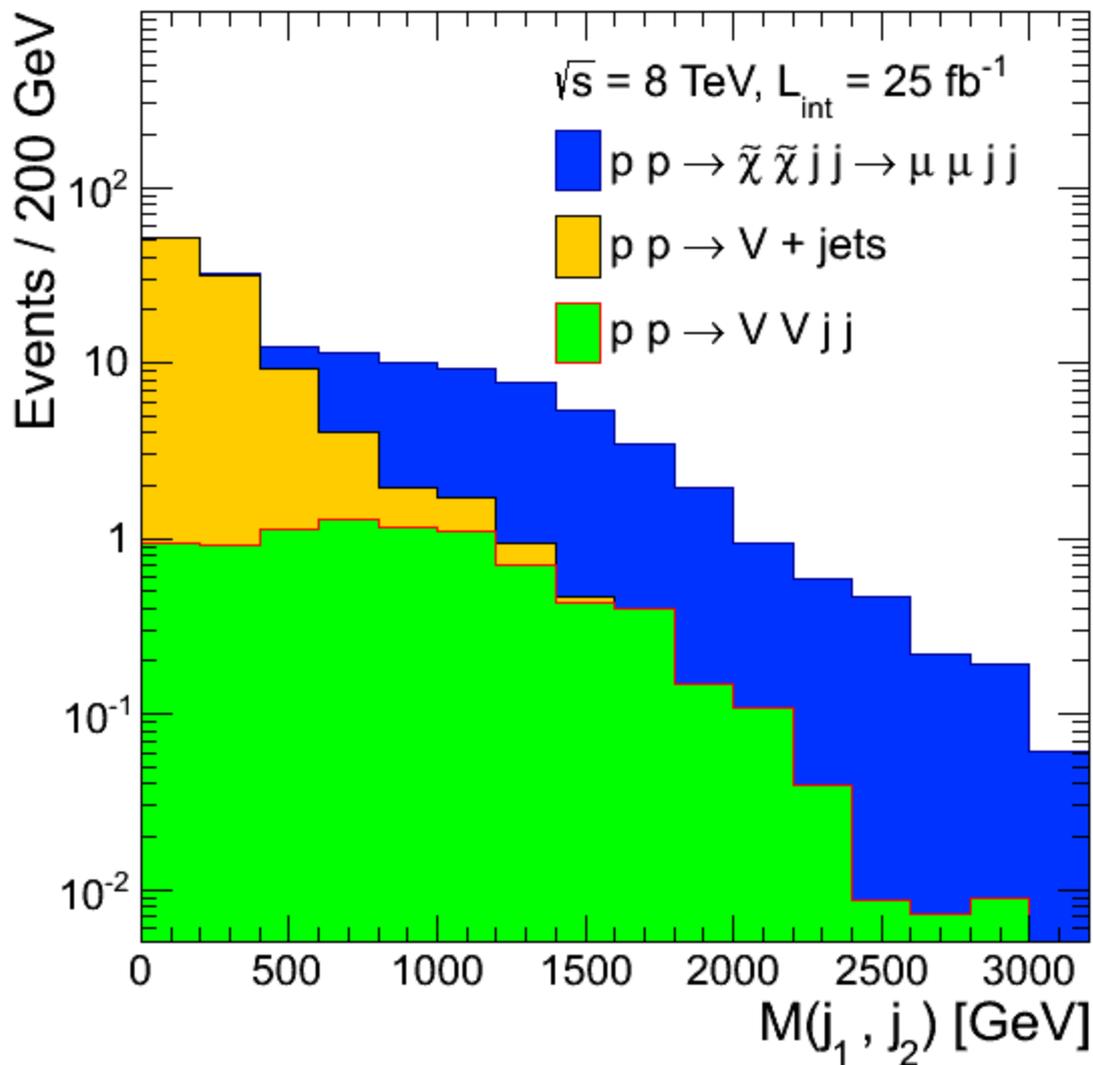
$$\sqrt{s} = 8 \text{ TeV}$$

Lum: 25 fb⁻¹

Two isolated μ 's with $p_T > 20$ GeV in $\eta < 2.1$

$$\text{For } 3\sigma : m_{\tilde{\chi}_1^\pm} \sim m_{\tilde{\chi}_2^0} = 330 \text{ GeV}$$

Signal: $\geq 2j+2\mu$ +missing energy



Conclusion

Annihilation diagrams: mostly non-colored particles, e.g., sleptons, staus, charginos, neutralinos, etc

→ Investigate sleptons, charginos, neutralinos etc. at the LHC

- **Sleptons, charginos etc. can be produced via cascade decays: squarks, gluinos etc**

Use the signatures and BEST to construct a decision tree

→ determine model parameters and the relic density based on the LHC measurements

- **Sleptons, charginos etc. can be produced via vector boson fusion**

Use high E_T forward jets in opposite hemispheres with large dijet invariant mass