

Baryon Asymmetry, Dark Matter & Neutrino Mass via Exotic Multiplets

SANDY S. C. LAW

National Cheng Kung University
Tainan, Taiwan



Motivation & Overview



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Some related references: Cai et al., *JHEP* 1112 054 (2011) [arXiv:1108.0969 [hep-ph]]
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A Model with Exotic Multiplets



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We add to the SM the following **exotic $SU(2)_L$ multiplet** fields:-

fermion 5-plets

$N_k \sim (1, 5, 0)$

$\times 3$ generations ($k = 1, 2, 3.$)

scalar 6-plets

$\chi \sim (1, 6, -1/2)$

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A Model with Exotic Multiplets [2]



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the subscripts α , β denote the many independent ways to contract the components

$$\text{e.g. } \phi^\dagger \phi \chi^\dagger \chi : \underbrace{2^* \otimes 2}_{1+3} \otimes \underbrace{6^* \otimes 6}_{1+3+\dots} \quad \text{or} \quad \underbrace{2 \otimes 6}_{5+7} \otimes \underbrace{2^* \otimes 6^*}_{5^*+7^*} \quad \text{or} \quad \underbrace{2 \otimes 6^*}_{5^*+7^*} \otimes \underbrace{2^* \otimes 6}_{5+7} .$$

however, only a subset of these are truly independent.

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- $\chi^\dagger \chi \chi \phi \ll 1$ (NB: technically natural) but exactly how small depends on the situation
- $\mu_\phi, \mu_\chi, \lambda_\phi, \lambda_{\chi\alpha}, \dots$ must be such that $\text{VEV} \langle \chi \rangle = 0$

A Model with Exotic Multiplets [3]



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- When $\chi^\dagger \chi \chi \phi \ll 1$ and $\langle \chi \rangle = 0$, the Lagrangian is also invariant under

$$\Psi_{\text{SM}} \rightarrow \Psi_{\text{SM}} ; N_k \rightarrow -N_k ; \chi \rightarrow -\chi$$

which ensures the lightest fermion 5-plet N_k (e.g. N_1) be absolutely stable if $M_\chi > M_1$.

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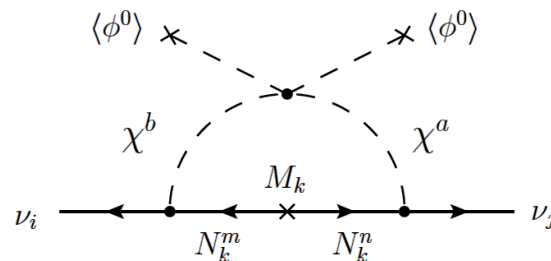
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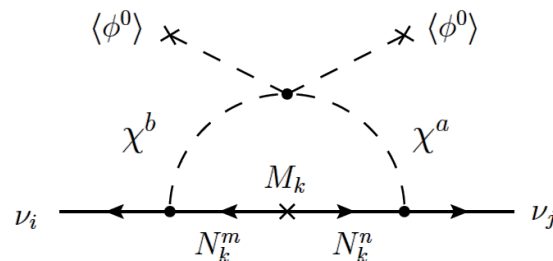
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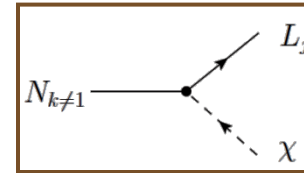


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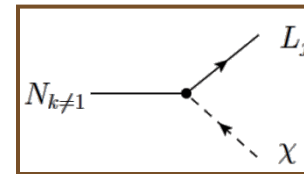


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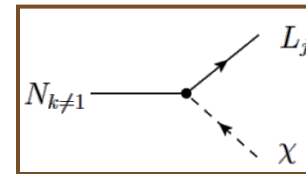


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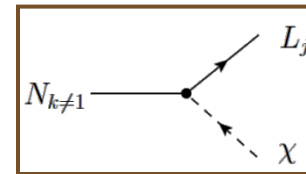


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baryon asymmetry

- So, the challenge is to demonstrate that there exists a **parameter space** where all three problems can be addressed consistently.

A Model with Exotic Multiplets [5]



The **key parameters** in the model at a glance:

$$\mathbf{M}_1 \quad , \quad \mathbf{M}_2 \quad , \quad \mathbf{M}_3 \quad , \quad \mathbf{M}_\chi \quad , \quad h_{jk} \quad , \quad \lambda'_{\phi\chi}$$

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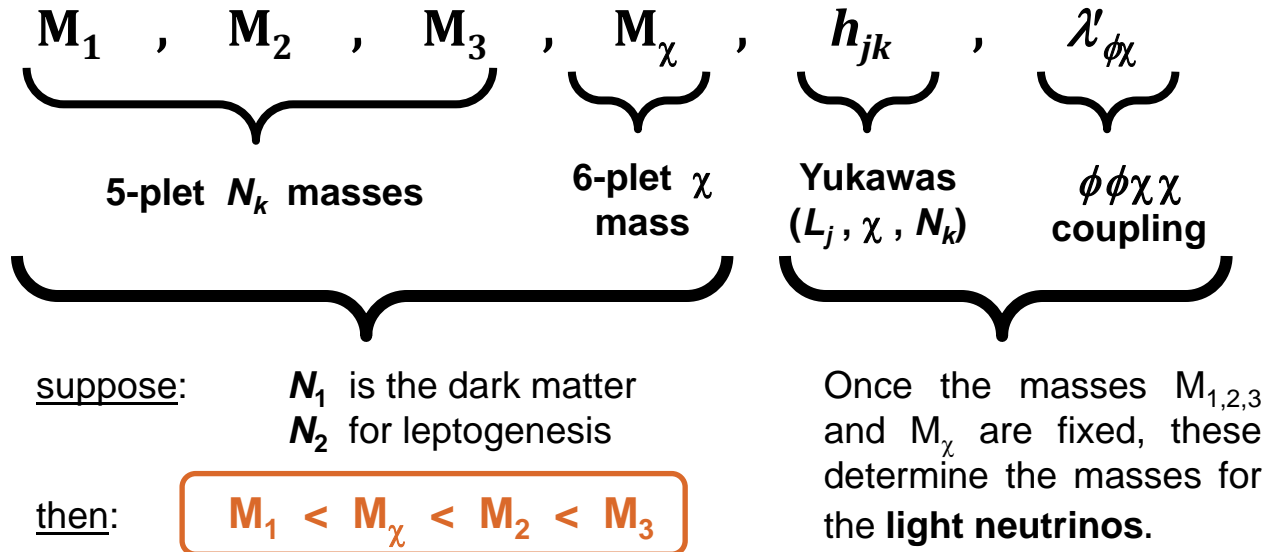
suppose: N_1 is the dark matter
 N_2 for leptogenesis

then: $M_1 < M_\chi < M_2 < M_3$

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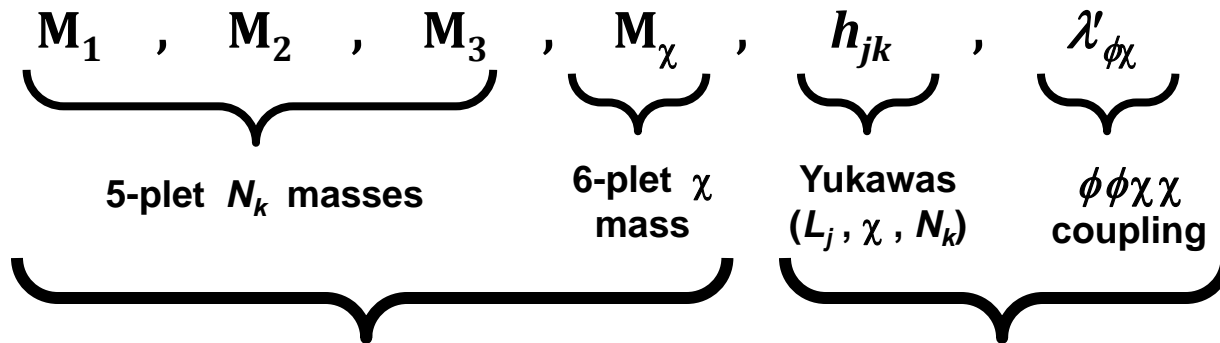
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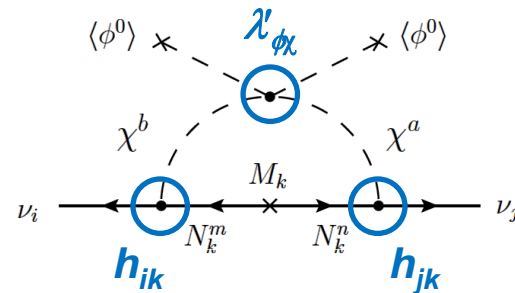
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suppose: N_1 is the dark matter
 N_2 for leptogenesis

then: $M_1 < M_\chi < M_2 < M_3$

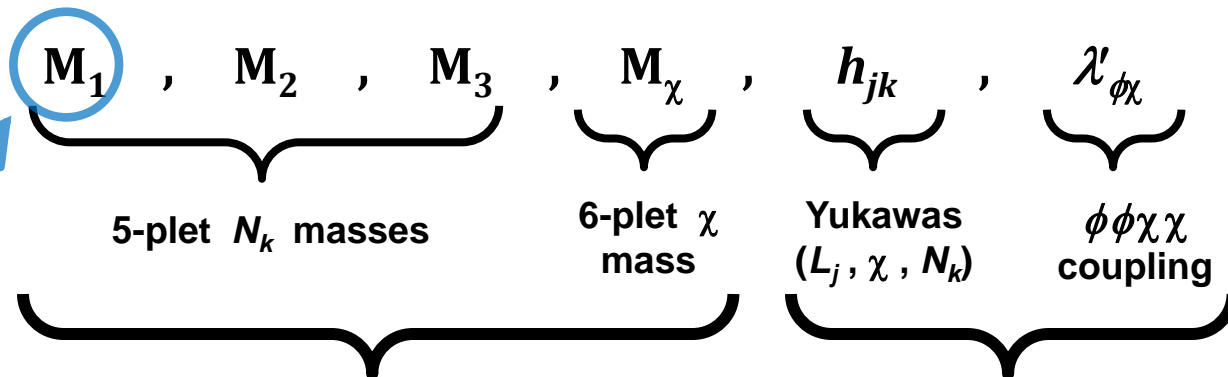
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A Model with Exotic Multiplets [5]



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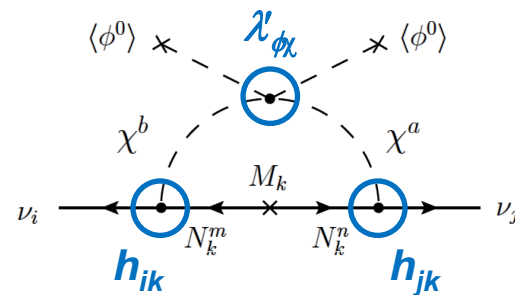
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The M_1 scale is dictated by the constraints from DM (e.g. **relic density**):

$$M_1 \gtrsim 10 \text{ TeV}$$

(co)annihilation of N_k 's mediated by SM gauge bosons assumed [Cirelli et al., New J. Phys. 11,105005 (09)]

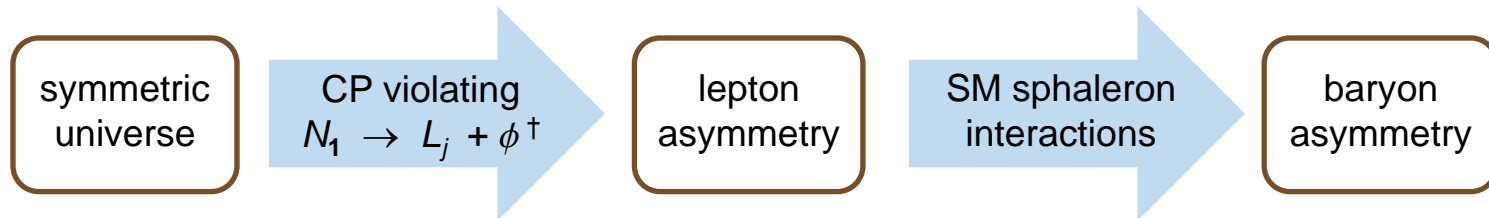


Baryogenesis via N_2 -Leptogenesis



Baryogenesis via N_2 -Leptogenesis

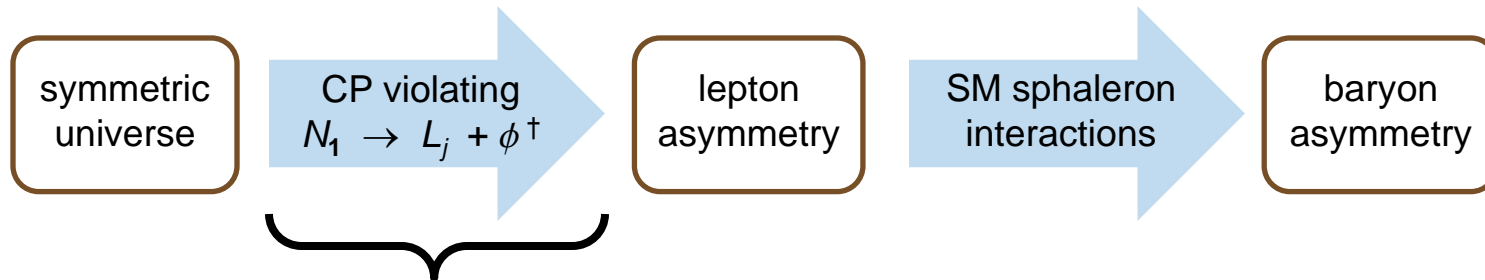
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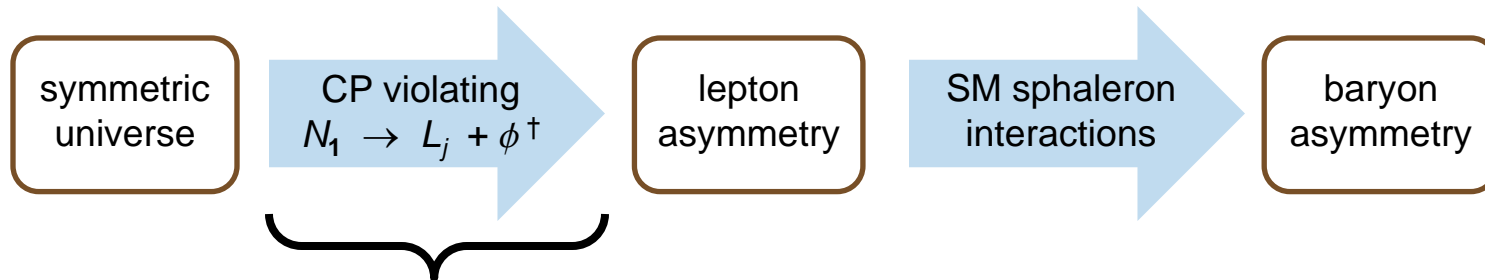


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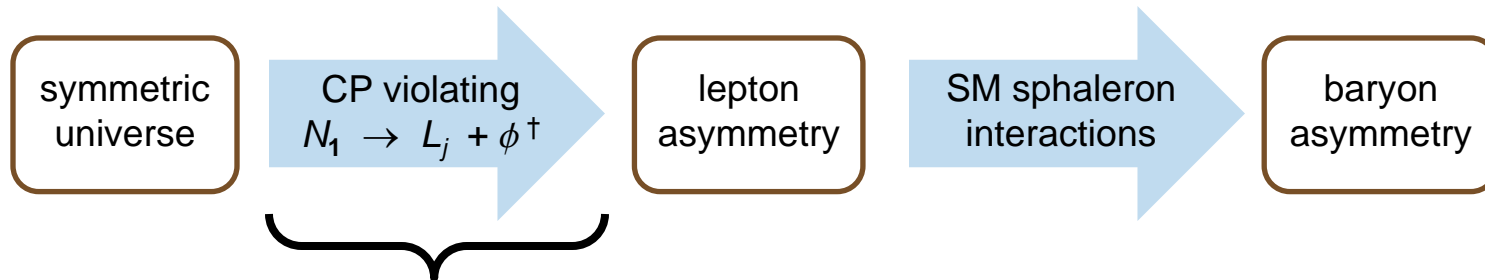


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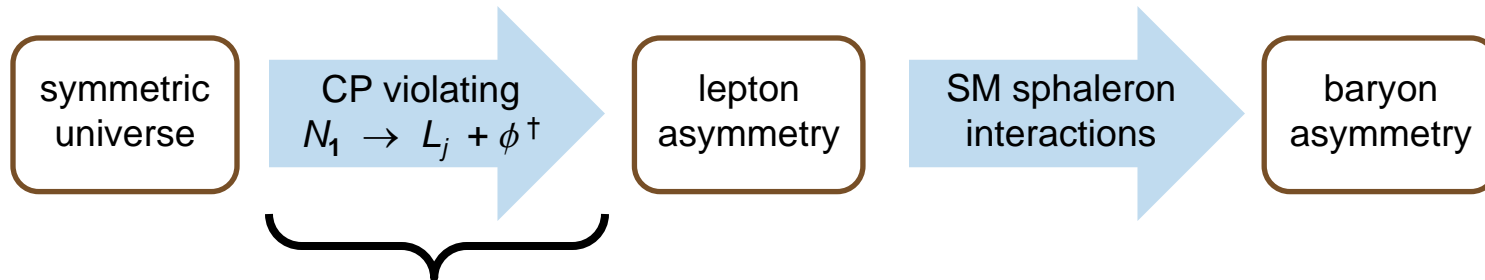


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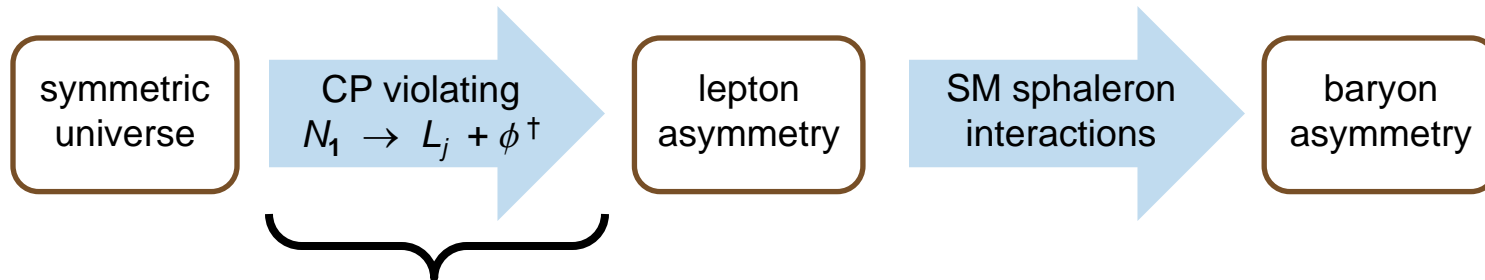


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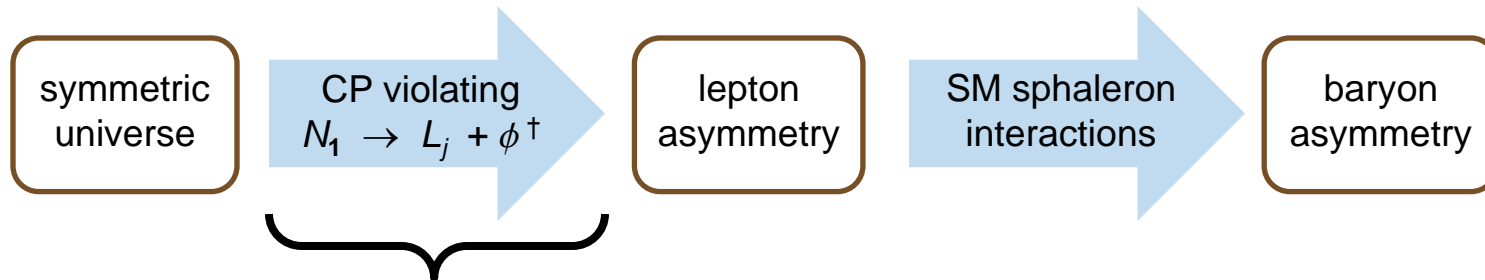


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For our case with $M_\chi > M_1$, $T \simeq M_\chi$ and $\chi \rightarrow L_j N_1$ is out-of-equilibrium.

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NB: the decay of χ will NOT generate any lepton asymmetry since there is only **one** type of 6-plet scalar in the model \Rightarrow vanishing absorptive part for the interference term with one-loop correction graphs.

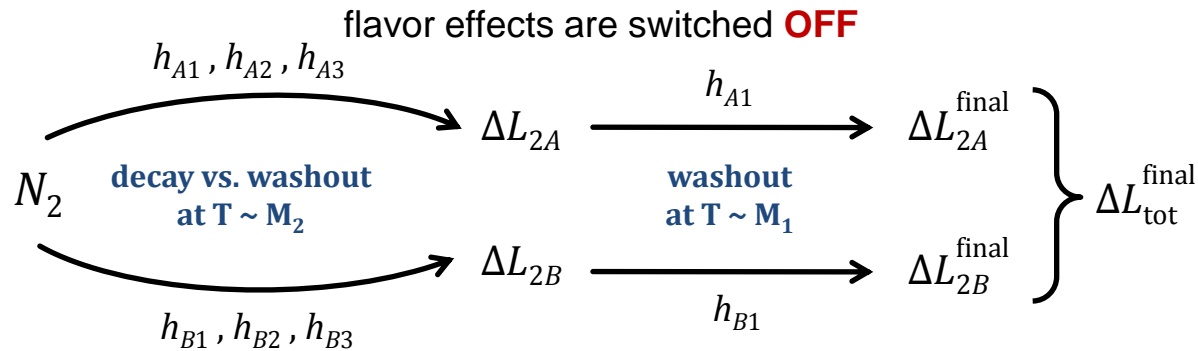
Flavored vs. Un-flavored N_2 -Leptogenesis



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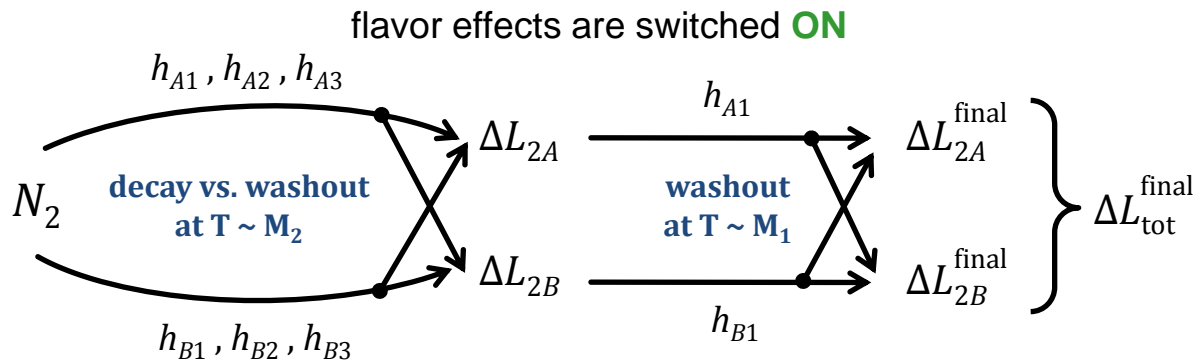
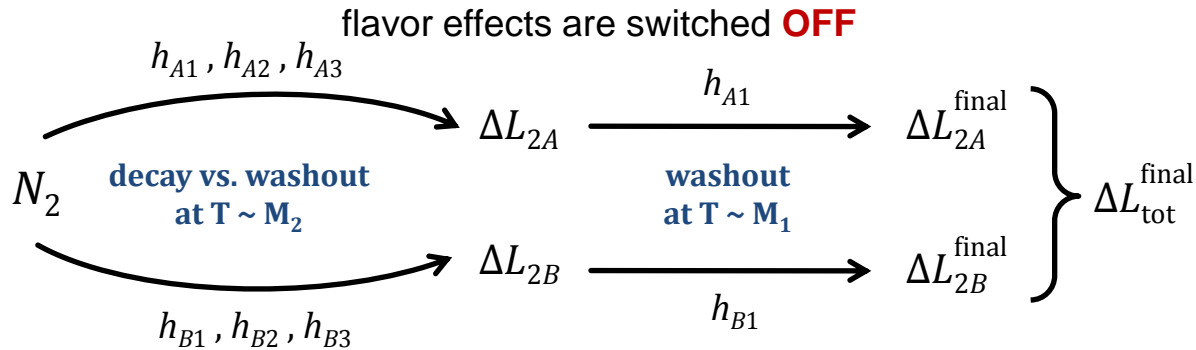
Schematic visualization of how flavor effects can make a difference



Flavored vs. Un-flavored N_2 -Leptogenesis



Schematic visualization of how flavor effects can make a difference



The interdependence of parameters is less restrictive when flavor effects is **ON**. So, it becomes **possible** to find a set that can generate enough ΔL .

5-plet N_2 -Leptogenesis with flavors



Recall that we require the mass relation:-

$$M_1 < M_\chi < M_2 < M_3$$

5-plet N_2 -Leptogenesis with flavors



Let's assume further that the spectrum is hierarchical:-

$$M_1 \ll M_\chi \ll M_2 \ll M_3$$

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$$\begin{aligned} \frac{d\mathcal{N}_{N_2}}{dz} &= -D_2 (\mathcal{N}_{N_2} - \mathcal{N}_{N_2}^{\text{eq}}) , \\ \frac{d\mathcal{N}_{\Delta_\perp}}{dz} &= -\varepsilon_{2\perp} D_2 (\mathcal{N}_{N_2} - \mathcal{N}_{N_2}^{\text{eq}}) - P_{2\perp}^0 W_2 \sum_{j=\perp, \tau} C_{\perp j}^{f=2} \mathcal{N}_{\Delta_\perp} , \\ \frac{d\mathcal{N}_{\Delta_\tau}}{dz} &= -\varepsilon_{2\tau} D_2 (\mathcal{N}_{N_2} - \mathcal{N}_{N_2}^{\text{eq}}) - P_{2\tau}^0 W_2 \sum_{j=\perp, \tau} C_{\tau j}^{f=2} \mathcal{N}_{\Delta_\tau} , \end{aligned}$$

where $z = M_2/T$.

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CP asymmetry
from N_2 decays

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decay term

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washout term
(dominated by inverse decays)

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tree-level
flavor projector

$$P_{2j}^0 = \frac{h_{j2}^* h_{j2}}{(h^\dagger h)_{22}}$$

5-plet N_2 -Leptogenesis with flavors



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non-diagonal flavor coupling matrix
(this leads to a coupled system)

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Similarly, during the **additional washout stage** (at $T \sim M_\chi$) for **three** flavors.

5-plet N_2 -Leptogenesis with flavors



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observe that for this stage, there is no contribution from decays.

Some comments on the solution



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- The resultant set of Yukawa's, h_{jk} , must be checked for consistency with known **neutrino data** by calculating their effects on the 1-loop neutrino diagram.

A workable example (normal hierarchy)



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Recalling the key parameters in the theory:

$$\mathbf{M}_1, \mathbf{M}_\chi, \mathbf{M}_2, \mathbf{M}_3, h_{jk}, \lambda'_{\phi\chi}$$

$\mathcal{O}(10^4) \text{ GeV}$

fixed by DM
constraints

A workable example (normal hierarchy)



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M_1	,	M_χ	,	M_2	,	M_3	,	h_{jk}	,	$\lambda'_{\phi\chi}$
$O(10^4)\text{ GeV}$		10^7		10^{10}		10^{13}				
fixed by DM constraints				GeV (hierarchical spectrum)						

M_1 , M_χ , M_2 , M_3 , h_{jk} , $\lambda'_{\phi\chi}$

$O(10^4) \text{ GeV}$ 10^7 10^{10} GeV 10^{13}

fixed by DM constraints (hierarchical spectrum)

The diagram shows six mass parameters arranged horizontally: M_1 , M_χ , M_2 , M_3 , h_{jk} , and $\lambda'_{\phi\chi}$. Below M_1 is the value $O(10^4) \text{ GeV}$ followed by the text "fixed by DM constraints". Below M_χ is the value 10^7 . Below M_2 is the value 10^{10} GeV followed by "(hierarchical spectrum)". Below M_3 is the value 10^{13} . The parameter h_{jk} is circled in blue, and a blue curved arrow points from the bottom of the circle down towards the "(hierarchical spectrum)" label.

For illustration, we use $K_{2\tau} \simeq 65$ and $K_{\chi\tau} \simeq 0.1$

A workable example (normal hierarchy)



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 fixed by DM constraints GeV (hierarchical spectrum)

The requirement for $K_{2\tau} \gtrsim 1$ and $K_{\chi\tau} \ll 1$ will constrain the Yukawa $h_{\tau k}$ for all k .

For illustration, we use $K_{2\tau} \simeq 65$ and $K_{\chi\tau} \simeq 0.1$

May tune this to control the size of the Yukawa couplings given a light neutrino mass scale

A workable example (normal hierarchy)



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May tune this to control the size of the Yukawa couplings given a light neutrino mass scale

With these parameters and assuming a **normal hierarchy** of light neutrinos (with e.g. $m_1 \simeq 0.002 \text{ eV}$), a possible set of h_{jk} that is consistent with oscillation data is

A workable example (normal hierarchy)



Recalling the key parameters in the theory:

$$\begin{array}{ccccccc}
 \mathbf{M}_1 & , & \mathbf{M}_\chi & , & \mathbf{M}_2 & , & \mathbf{M}_3 & , & \mathbf{h}_{jk} & , & \lambda'_{\phi\chi} \\
 \mathbf{O(10^4) GeV} & & \mathbf{10^7} & & \mathbf{10^{10}} & & \mathbf{10^{13}} & & & & \mathbf{0.1} \\
 \text{fixed by DM} & & & & \mathbf{GeV} & & & & & & \uparrow \\
 \text{constraints} & & & & \mathbf{(hierarchical spectrum)} & & & & & &
 \end{array}$$

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$$\left. \begin{array}{l}
 h_{e1} = 1.23 + 0.359i, \quad h_{e2} = 0.104 - 0.329i, \quad h_{e3} = -0.344 + 0.263i, \\
 h_{\mu1} = 1.71 - 1.02i, \quad h_{\mu2} = -0.304 - 0.468i, \quad h_{\mu3} = -3.76 + 0.367i, \\
 h_{\tau1} = 1.07 \times 10^{-5}, \quad h_{\tau2} = 8.88 \times 10^{-3}, \quad h_{\tau3} = 5.34.
 \end{array} \right\} \eta_B \simeq 6 \times 10^{-10}$$

Other examples (inverted & quasi-degenerate)



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For the **inverted hierarchical** light neutrino spectrum and with $m_1 \simeq 0.002$ eV

$\lambda'_{\phi\chi} = 1$ to keep Yukawa's size perturbative, while $K_{z\tau} \simeq 2$, $K_{x\tau} \simeq 0.01$.

$$\begin{aligned} h_{e1} &= 2.73 - 2.63i, & h_{e2} &= -0.737 - 0.758i, & h_{e3} &= 0.592 + 0.353i, \\ h_{\mu1} &= 0.351 + 1.17i, & h_{\mu2} &= 0.329 - 0.098i, & h_{\mu3} &= 1.29 + 0.045i, \\ h_{\tau1} &= 3.40 \times 10^{-6}, & h_{\tau2} &= 1.56 \times 10^{-3}, & h_{\tau3} &= 1.61. \end{aligned}$$

These will give $\eta_B \simeq 1.5 \times 10^{-10}$ which is marginally successful.

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For the **quasi-degenerate** light neutrino spectrum and with $m_1 \simeq 0.16$ eV

$$\lambda'_{\phi\chi} = 1 \quad \text{and} \quad K_{z\tau} \simeq 1.2, \quad K_{x\tau} \simeq 0.01$$

$$\begin{aligned} h_{e1} &= 3.25 - 1.91i, & h_{e2} &= 0.541 + 0.895i, & h_{e3} &= -0.090 - 0.128i, \\ h_{\mu1} &= 1.972 + 3.23i, & h_{\mu2} &= -0.916 + 0.543i, & h_{\mu3} &= 0.040 - 0.078i, \\ h_{\tau1} &= 3.40 \times 10^{-6}, & h_{\tau2} &= 1.21 \times 10^{-3}, & h_{\tau3} &= 4.06. \end{aligned}$$

These will give $\eta_B \simeq 4.3 \times 10^{-11}$ which is the least favored case.

Summary



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- We have demonstrated that there is a parameter space in this model where a consistent solution to all three problems can be obtained.